

Search for Signatures of Extra Dimensions in the Diphoton Mass Spectrum at the Large Hadron Collider

S. Chatrchyan *et al.**

(CMS Collaboration)

(Received 4 December 2011; published 12 March 2012)

A search for signatures of extra spatial dimensions in the diphoton invariant-mass spectrum has been performed with the CMS detector at the LHC. No excess of events above the standard model expectation is observed using a data sample collected in proton-proton collisions at $\sqrt{s} = 7$ TeV corresponding to an integrated luminosity of 2.2 fb^{-1} . In the context of the large-extra-dimensions model, lower limits are set on the effective Planck scale in the range of 2.3–3.8 TeV at the 95% confidence level. These limits are the most restrictive bounds on virtual-graviton exchange to date. The most restrictive lower limits to date are also set on the mass of the first graviton excitation in the Randall-Sundrum model in the range of 0.86–1.84 TeV, for values of the associated coupling parameter between 0.01 and 0.10.

DOI: [10.1103/PhysRevLett.108.111801](https://doi.org/10.1103/PhysRevLett.108.111801)

PACS numbers: 13.85.Rm, 11.25.Wx, 13.85.Qk

Over a decade ago, Arkani-Hamed, Dimopoulos, and Dvali (ADD) [1,2] proposed that extra spatial dimensions could potentially solve the standard model (SM) hierarchy problem [3], which consists of the observation of the unnatural difference of scales between the gravitational and electroweak theories. They proposed a scenario whereby the SM is constrained to the common $3 + 1$ space-time dimensions (brane), while gravity is free to propagate throughout a larger multidimensional space (bulk). The gravitational flux on the brane is therefore diluted by virtue of Gauss's law in the bulk, which relates the fundamental Planck scale M_D to the apparent reduced scale $\bar{M}_{\text{Pl}} \approx 2 \times 10^{18} \text{ GeV}$ according to the formula $M_D^{n_{\text{ED}}+2} = \frac{\bar{M}_{\text{Pl}}^2}{r^{n_{\text{ED}}}}$, where r and n_{ED} are the size and number of the extra dimensions (ED), respectively. Postulating a fundamental Planck scale to be on the order of the electroweak symmetry breaking scale (~ 1 TeV) results in ED with $r < 1$ mm for $n_{\text{ED}} \geq 2$.

Another model of ED that solves the hierarchy problem was proposed by Randall and Sundrum (RS) [4]. In this model, the observed hierarchy is related instead to the warped anti-de Sitter (AdS) geometry of the ED. We specifically consider the RS1 model whereby only one finite ED exists separating two branes, one at each end of the ED. The geometry of the bulk is based on a slice of a five-dimensional AdS space with a length πr_c , where r_c is the compactification radius. The full metric is given by $ds^2 = e^{-kr_c y} \eta_{\mu\nu} dx^\mu dx^\nu - r_c^2 dy^2$, where Greek indices run over four-dimensional space-time, $\eta_{\mu\nu}$ is the

Minkowski metric tensor, and $0 \leq y \leq \pi$ is the coordinate along the single ED of radius r_c . The value of k specifies the curvature scale (or “warp factor”) and relates the fundamental Planck scale on one brane to the apparent scale on the other by $\Lambda_\pi = \bar{M}_{\text{Pl}} e^{-kr_c \pi}$. As a consequence, a value of $kr_c \sim 10$ would provide a natural solution to the hierarchy problem, yielding $\Lambda_\pi \sim 1$ TeV.

Phenomenologically, the excited gravitons in both models preferentially decay into two gauge bosons, such as photons, rather than into two leptons, because the graviton has spin 2, and so fermions cannot be produced in an s wave. In the RS scenario, gravitons appear as well-separated Kaluza-Klein (KK) excitations with masses and widths determined by the parameters of the RS1 model. One convenient choice of parametrization is the mass M_1 of the first excitation of the graviton and the dimensionless warp factor $\tilde{k} \equiv k/\bar{M}_{\text{Pl}}$, which defines the strength of associated coupling to the SM fields. Precision electroweak data constrain $\tilde{k} \geq 0.01$, while perturbativity requirements limit $\tilde{k} \leq 0.10$ [5].

In the ADD model, the wave function of the KK gravitons must satisfy periodic boundary conditions, resulting in discrete energy levels with modal spacing of the order of the inverse ED size, from 1 to 100 meV, much smaller than the spacing in the RS1 model, which is expected to be of the order of 1 TeV. This effect produces an apparent continuum spectrum of diphotons, rather than distinct resonances, at high (~ 1 TeV) diphoton invariant mass $M_{\gamma\gamma}$.

Summing over all KK modes in the ADD scenario results in a divergence in the cross section, so an ultraviolet (UV) cutoff scale M_ζ is imposed. This effective Planck scale is related to—but potentially different from—the fundamental Planck scale M_D . The precise relationship depends on the UV completion of the effective theory. The effects of virtual-graviton production on the differential diphoton cross section are parametrized by the single

*Full author list given at the end of the article.

Published by the American Physical Society under the terms of the [Creative Commons Attribution 3.0 License](https://creativecommons.org/licenses/by/3.0/). Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI.

variable $\eta_G = \mathcal{F}/M_S^4$, where \mathcal{F} is an order-unity dimensionless parameter, for which several conventions exist [6]. In this Letter, we set lower limits on M_S in three different conventions: GRW [7], Hewett [8], and HLZ [9].

Searches for ED via virtual-graviton effects in the ADD model have been conducted at HERA, LEP, the Tevatron, and the LHC [10,11]. The most stringent previously published limits on M_S for $n_{\text{ED}} \geq 3$ come from the previous measurement in the diphoton channel at the Compact Muon Solenoid (CMS) experiment [6]. For $n_{\text{ED}} = 2$, measurements by the D0 experiment in the dijet [12] and diphoton + dielectron [13] channels are more restrictive. The most sensitive previous search for RS gravitons was conducted by the ATLAS experiment [14]. They used a search in the dilepton final state to exclude $M_1 < 1.63$ TeV for $\tilde{k} = 0.10$.

In this Letter, we present a search for both nonresonant and resonant diphoton production, in the ADD and RS1 models, respectively. We use data corresponding to an integrated luminosity of 2.2 fb^{-1} , collected in pp collisions at $\sqrt{s} = 7$ TeV at the LHC with the CMS detector between March and August 2011.

The CMS detector [15] is designed to study collisions at the LHC. An all-silicon tracker, an electromagnetic calorimeter (ECAL), and a hadronic sampling calorimeter are all contained within a large-bore 3.8 T superconducting solenoid. In the central region, the tracker consists of three radial layers of silicon pixel detectors followed radially by silicon strip detectors. The finely segmented ECAL has a design resolution for unconverted photons better than 0.5% at energies exceeding 100 GeV in the barrel ($|\eta| < 1.44$). Here, the pseudorapidity η is defined as $-\ln[\tan(\theta/2)]$, where θ is the polar angle with respect to the direction of the counterclockwise beam. Beyond the solenoid lie four layers of muon detectors. The instantaneous luminosity is measured with a relative uncertainty of 4.5% using information from forward hadronic calorimeters [16]. The two-tiered CMS trigger consists of the level-one trigger, composed of custom hardware, and the software-based high-level trigger.

Events for the analysis were collected through a diphoton trigger, where each photon was required to have a transverse energy $E_T \equiv E \sin\theta$ of at least 33, 55, or 60 GeV, depending on the instantaneous luminosity. We require that an event be consistent with a pp collision and have at least one well-reconstructed primary vertex [17]. We then reconstruct photons with $E_T > 70$ GeV in the ECAL barrel by clustering electromagnetic energy depositions. Electrons that do not originate from photon conversions are suppressed by using information from the pixel detector to associate tracks and ECAL clusters compatible with an electron hypothesis. The probability of misidentification of an electron as a photon is approximately 3%, resulting in a negligibly small contribution to the diphoton spectrum in the signal region.

Hadronic jets can be misidentified as photons when their leading hadron is an energetic π^0 or η meson. We reduce the misidentification rate from this source by placing the same restrictions on the isolation as in the previous analysis for this channel [6]. These restrictions limit the total transverse energy because of tracks and calorimeter depositions near the photon cluster. Restrictions on the shower-shape variable $\sigma_{\eta\eta}$, which is a modified second moment of the electromagnetic energy cluster about its mean η position [18], also suppress hadronic misidentification. Topological and timing criteria suppress anomalous signals present in the ECAL [19]. Diphoton events are selected in which $M_{\gamma\gamma} > 140$ GeV.

The photon reconstruction and identification efficiency is determined in Monte Carlo (MC) simulation and corrected using a data-to-MC scale factor of 1.005 ± 0.034 derived from studying $Z \rightarrow e^+e^-$ events. The measured efficiency for a single $E_T > 70$ GeV photon with $|\eta| < 1.44$ is $(87.4 \pm 5.4)\%$ and depends only weakly on the E_T and η of the photon, and the number of extra collisions present in the event. The systematic uncertainty bounds the variation as a function of these variables, the most significant of which is the number of extra collisions. We reweight the simulation to give the same reconstructed primary-vertex distribution (on average 6–8 vertices) as observed in the data. We determine the corresponding diphoton reconstruction and identification efficiency $(76.4 \pm 9.6)\%$ by squaring the single-photon efficiency.

The simulation of ED in the ADD model is performed using version 1.3.0 of the SHERPA [20] MC generator. The simulation includes both SM diphoton production and signal diphoton production via virtual-graviton exchange in order to account for the interference effects between the SM and ADD processes. The leading-order (LO) SHERPA cross sections are multiplied by a constant next-to-leading-order (NLO) K factor of 1.6 ± 0.1 , a value that represents an updated calculation for $\sqrt{s} = 7$ TeV by the authors of Refs. [21,22]. The systematic uncertainty on the signal K factor reflects the approximate variation of the K factor over a large region of the model parameters; it is not intended to account for the theoretical uncertainty. The cross sections in the simulation are conservatively set to zero for $\sqrt{\hat{s}} > M_S$ because the theory becomes nonperturbative for larger values of $\sqrt{\hat{s}}$. Introducing this sharp truncation reduces the upper limits on M_S by a few percent.

The simulation of RS-graviton production is performed using version 6.424 of the PYTHIA [23] MC program. The signal cross section is scaled by a mass-dependent NLO K factor [21,22], which ranges from 1.6 to 1.8 as a function of $M_{\gamma\gamma}$ and for different values of \tilde{k} . The CTEQ6L1 [24] parton distribution functions (PDF) are used in the simulation of both the ADD and RS models, and a 1.5% relative uncertainty on the signal acceptance is included by measuring its dependence on the choice of PDF and its uncertainties.

Optimization of the event selection is done separately for both ADD and RS scenarios. The signal in both cases is predominantly at central values of η , while the high- $M_{\gamma\gamma}$ SM background dominates the signal in the forward region; therefore, we restrict ourselves to photons located in the ECAL barrel only. In the ADD scenario, we find that the optimal region for the search, based on the expected signal significance, is $M_{\gamma\gamma} > 900$ GeV. This choice of selection depends weakly on the model parameters.

In the search for RS gravitons, a fixed window is selected about the M_1 mass point of interest. Because the signal shapes deviate from Gaussian distributions, we define an effective measure of the signal width σ_{eff} as the half-width of the narrowest mass interval containing 68% of the signal from simulation. The value of σ_{eff} ranges from 6 to 21 GeV for RS gravitons with M_1 between 500 and 2000 GeV and $\tilde{k} = 0.01$. The dependence on M_1 is linear and also increases with \tilde{k} ($\sigma_{\text{eff}} = 42$ GeV for $M_1 = 2$ TeV and $\tilde{k} = 0.10$). A window is then formed about the resonance mean of size $\pm 5\sigma_{\text{eff}}$ in the data. This window contains 96%–97% of the signal acceptance for all mass points considered in this analysis, and the detector resolution is negligible with respect to the window size. This choice of the window maximizes the signal acceptance and analysis sensitivity in the case of small backgrounds.

Backgrounds from the misidentification of a hadronic jet as a photon are small in the signal region but contribute to the low- $M_{\gamma\gamma}$ region. Two such sources of backgrounds from isolated-photon misidentification are considered: multijet production and prompt single-photon ($\gamma + \text{jet}$) production. In particular, we measure on a background-dominated sample a misidentification rate, defined as the ratio of the number of isolated photon candidates to non-isolated photonlike objects. These photonlike objects are reconstructed as photons but fail one of the isolation or shower-shape criteria; therefore, the samples corresponding to numerator and denominator are mutually exclusive, and prompt photons have a negligibly small contribution to the denominator. The misidentification rate is measured in a photon-triggered sample in bins of photon(like) candidate E_T , but the objects used in the measurement are required to be well separated from the triggered object to avoid a trigger-induced bias.

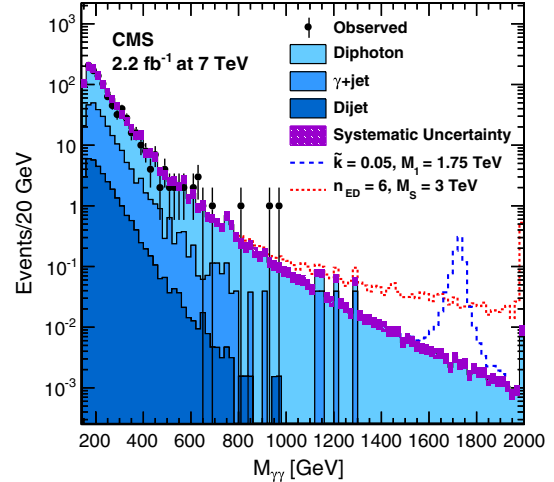


FIG. 1 (color online). Observed event yields (points with error bars) and background expectations (filled solid histograms) as a function of the diphoton invariant mass. Photons are required to be isolated, with $E_T > 70$ GeV and $|\eta| < 1.44$. The shaded band around the background estimation corresponds to the average systematic uncertainty over the spectrum. The precise per-bin uncertainty is not provided for the sake of clarity. The last bin includes the sum of all contributions for $M_{\gamma\gamma} > 2.0$ TeV. The simulated distributions for two, nonexcluded signal hypotheses are shown for comparison as dotted (ADD) and dashed (RS) lines.

Because the background-dominated sample in which we measure the misidentification rate may contain some genuine, isolated photons that “contaminate” the numerator of the misidentification rate, we correct for this on a bin-by-bin basis. The $\sigma_{\eta\eta}$ requirement is released and the numerator sample is fit for the fraction of prompt photons using one-dimensional probability density histograms (“templates”) in $\sigma_{\eta\eta}$. The signal template is constructed from MC simulation, and the background template is constructed from reconstructed photons that fail one or more of the isolation criteria. The measured misidentification rate falls from 7% at $E_T = 70$ GeV to 2% at $E_T = 120$ GeV. We apply a 20% systematic uncertainty to the rate derived from the variation of the misidentification rate measured in a jet-triggered sample.

TABLE I. Observed event yields and background expectations for different reconstructed diphoton invariant-mass ranges. Full systematic uncertainties are included.

Process	Diphoton invariant-mass range [TeV]			
	[0.14, 0.2]	[0.2, 0.5]	[0.5, 0.9]	>0.9
Multijet	15 ± 6	17 ± 7	0.2 ± 0.1	0.003 ± 0.001
$\gamma + \text{jet}$	102 ± 15	124 ± 18	2.5 ± 0.4	0.19 ± 0.04
Diphoton	372 ± 70	414 ± 78	16.9 ± 3.2	1.3 ± 0.3
Backgrounds	489 ± 73	555 ± 81	19.6 ± 3.2	1.5 ± 0.3
Observed	484	517	16	2

The multijet and $\gamma + \text{jet}$ backgrounds to the reconstructed diphoton spectrum are estimated by using the misidentification rate to extrapolate from two background-dominated reference regions, both selected with the same diphoton trigger as the primary signal sample. One region includes events with only one isolated photon, but one or more nonisolated photons. The other region includes events with no isolated photons, but two or more nonisolated photons. The diphoton trigger is sufficiently inclusive that the regions are unaffected by the trigger selection. By applying the prompt-photon misidentification rate to these two reference regions, we predict the $\gamma + \text{jet}$ and multijet backgrounds in the signal region.

The SM diphoton background dominates the signal region. The expected number of background events due to this process is computed by rescaling the prediction from PYTHIA with a NLO K factor that varies with $M_{\gamma\gamma}$. The NLO prediction is calculated with the DIPHOX+GAMMA2MC [25,26] generators, which take into account the fragmentation processes in which the photons can come from the collinear fragmentations of hard partons. A separate analysis by CMS has also demonstrated good agreement with the NLO prediction at low $M_{\gamma\gamma} \lesssim 300$ GeV [27]. The subleading-order gluon-fusion box diagram is included as a part of the PYTHIA calculation because of its large contribution at the LHC energy, although its effects are small at high $M_{\gamma\gamma}$. The K factor varies between 1.7 and 1.1 from low to high $M_{\gamma\gamma}$. A systematic uncertainty of 15% on the value of the K factor is determined by examining the PDF uncertainties and variation of the renormalization and factorization scales.

Figure 1 shows the invariant-mass distribution of the selected events, together with the estimated distributions for each of the backgrounds. Table I presents the observed number of events in the data and the predicted number of background events in different ranges in $M_{\gamma\gamma}$ and corresponds directly to Fig. 1. The last column corresponds to the signal region for the ADD search. We find that the observed data are consistent with the background estimate throughout the $M_{\gamma\gamma}$ spectrum and do not show an excess of events, neither resonant nor nonresonant.

TABLE II. The 95% C.L. lower limits on M_S (in TeV) in the GRW, Hewett, and HLZ conventions for two values of the ADD signal K factor, 1.0 and 1.6 ± 0.1 . All limits are computed with a signal cross section truncated to zero for $\sqrt{\hat{s}} > M_S$, where $\sqrt{\hat{s}}$ is the center-of-mass of the partonic collision. The limits are presented for both positive and negative interference in the Hewett convention and for $n_{\text{ED}} = 2-7$ in the HLZ convention. The median expected lower limits are given in parentheses.

K	Hewett			HLZ					
	GRW	Positive	Negative	$n_{\text{ED}} = 2$	$n_{\text{ED}} = 3$	$n_{\text{ED}} = 4$	$n_{\text{ED}} = 5$	$n_{\text{ED}} = 6$	$n_{\text{ED}} = 7$
1.0	2.94 (2.99)	2.63 (2.67)	2.28 (2.31)	3.29 (3.37)	3.50 (3.56)	2.94 (2.99)	2.66 (2.71)	2.47 (2.52)	2.34 (2.38)
1.6 ± 0.1	3.18 (3.24)	2.84 (2.90)	2.41 (2.44)	3.68 (3.77)	3.79 (3.85)	3.18 (3.24)	2.88 (2.93)	2.68 (2.73)	2.53 (2.58)

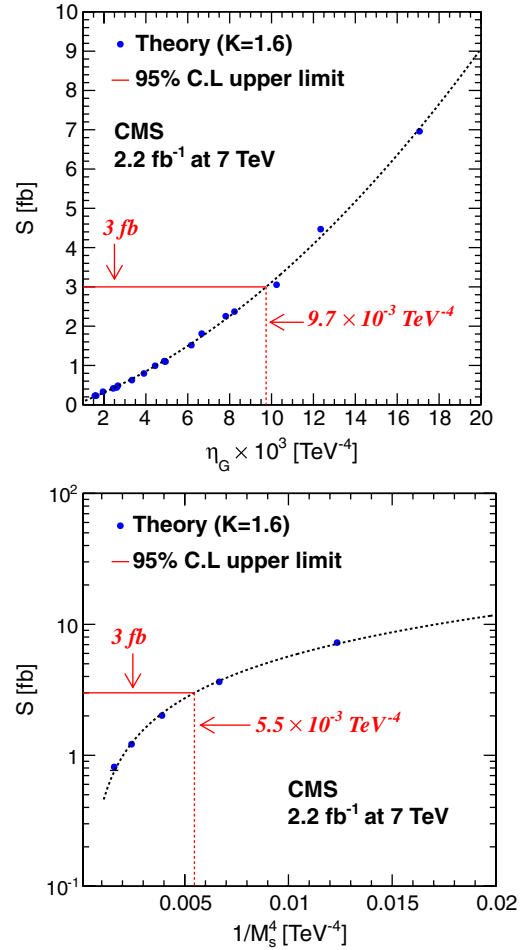


FIG. 2 (color online). Signal cross section S parametrization as a function of the strength of the ED effects η_G (top) and as a function of $1/M_S^4$ for the HLZ $n_{\text{ED}} = 2$ case (bottom).

To set limits on virtual-graviton exchange in the ADD scenario, we compare the number of observed and expected events in the signal region ($M_{\gamma\gamma} > 0.9$ TeV) and set 95% confidence level (C.L.) upper limits on the quantity $S \equiv (\sigma_{\text{total}} - \sigma_{\text{SM}}) \times \mathcal{B} \times \mathcal{A}$, where σ_{total} represents the total diphoton production cross section (including signal, SM, and interference effects), and σ_{SM} represents

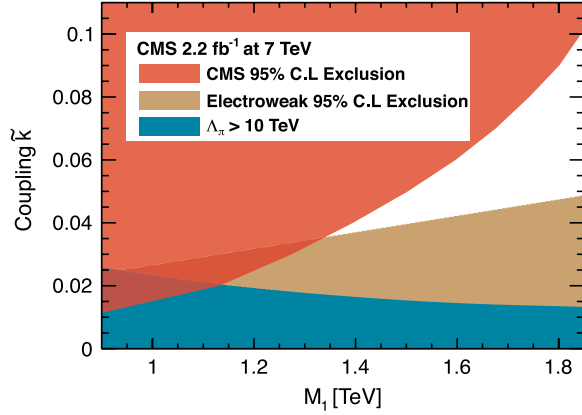


FIG. 3 (color online). The 95% C.L. exclusion region for the RS1 graviton model in the M_1 - \tilde{k} plane. The expected limits coincide very closely with the measured limits and so are not shown in the figure. Also shown are bounds due to electroweak constraints and naturalness ($\Lambda_\pi > 10$ TeV). Perturbativity requirements bound $\tilde{k} \lesssim 0.10$.

the SM diphoton production cross section. The signal branching fraction to diphotons is indicated by \mathcal{B} and the signal acceptance by \mathcal{A} . We use the CL_s technique [28,29] to compute the limits with a likelihood constructed from the Poisson probability to observe N events, given S , the signal efficiency ($76.4 \pm 9.6\%$), the expected number of background events (1.5 ± 0.3), and the integrated luminosity $\mathcal{L} = (2.2 \pm 0.1) \text{ fb}^{-1}$ [16]. The variation of the K factor is included in the statistical analysis as an uncertainty on the signal yield.

The observed (median expected) 95% C.L. upper limit on S is 3.0 fb (2.7 fb). For the HLZ $n_{\text{ED}} = 2$ case, we parametrize S directly as a smooth function of $1/M_S^4$. For all other conventions, S is parametrized as a function of the parameter η_G , as in Ref. [6]. The observed 95% C.L. limit, together with the signal parametrization, is shown in Fig. 2. The intersection of the cross-section limit with the parametrized curve determines the 95% C.L. upper limit on the parameter η_G . As seen from the plot, these upper limits on S correspond to upper limits of $\eta_G \leq 0.0097 \text{ TeV}^{-4}$ and $1/M_S^4 \leq 0.0055 \text{ TeV}^{-4}$. The upper limits on η_G are equated to lower limits on M_S and are shown in Table II.

For the RS scenario, the same limit-setting calculation is performed, but in a bounded window in $M_{\gamma\gamma}$. Figure 3 shows the excluded regions in the M_1 - \tilde{k} plane. Also shown are bounds due to precision electroweak measurements and

to naturalness arguments [5]. Table III presents the 95% C.L. lower limits on the graviton mass M_1 for different values of \tilde{k} . The median expected lower limits coincide within a few GeV of the observed limits.

In summary, we have performed a search for extra spatial dimensions leading to enhanced resonant or nonresonant diphoton production in proton-proton collisions at a center-of-mass energy of 7 TeV at the LHC. Using a data sample corresponding to an integrated luminosity of 2.2 fb^{-1} recorded by the CMS experiment, we observe no excess in diphoton production above the rate predicted from SM background sources. Values of the effective Planck scale M_S less than 2.3–3.8 TeV are excluded at 95% C.L. for ADD models. We also exclude at 95% C.L. resonant graviton production in the RS1 model with values of M_1 less than 0.86–1.84 TeV depending on the normalized coupling strength \tilde{k} . We present limits on both the ADD and RS1 models of extra dimensions in the diphoton final state that extend those observed at the D0 experiment [13], as well as those set previously by the CMS [6] and ATLAS [14] experiments.

We thank M. C. Kumar, P. Mathews, V. Ravindran, and A. Tripathi for the calculation of NLO K factors used in this Letter. We wish to congratulate our colleagues in the CERN accelerator departments for the excellent performance of the LHC machine. We thank the technical and administrative staff at CERN and other CMS institutes, and acknowledge support from FMSR (Austria); FNRS and FWO (Belgium); CNPq, CAPES, FAPERJ, and FAPESP (Brazil); MES (Bulgaria); CERN; CAS, MoST, and NSFC (China); COLCIENCIAS (Colombia); MSES (Croatia); RPF (Cyprus); Academy of Sciences and NICPB (Estonia); Academy of Finland, MEC, and HIP (Finland); CEA and CNRS/IN2P3 (France); BMBF, DFG, and HGF (Germany); GSRT (Greece); OTKA and NKTH (Hungary); DAE and DST (India); IPM (Iran); SFI (Ireland); INFN (Italy); NRF and WCU (Korea); LAS (Lithuania); CINVESTAV, CONACYT, SEP, and UASLP-FAI (Mexico); MSI (New Zealand); PAEC (Pakistan); SCSR (Poland); FCT (Portugal); JINR (Armenia, Belarus, Georgia, Ukraine, Uzbekistan); MST, MAE, and RFBR (Russia); MSTD (Serbia); MICINN and CPAN (Spain); Swiss Funding Agencies (Switzerland); NSC (Taipei); TUBITAK and TAEK (Turkey); STFC (United Kingdom); DOE and NSF (U.S.).

TABLE III. The 95% C.L. lower limits on M_1 for given values of the coupling parameter \tilde{k} . For $\tilde{k} < 0.03$, masses above the presented limits are excluded by electroweak and naturalness constraints. The median expected lower limits are numerically the same for the presented precision except for the $\tilde{k} = 0.01$ case, for which the expected lower limit on M_1 is 0.84 TeV.

\tilde{k}	0.01	0.02	0.03	0.04	0.05	0.06	0.07	0.08	0.09	0.10
M_1 [TeV]	0.86	1.13	1.27	1.39	1.50	1.59	1.67	1.74	1.80	1.84

- [1] N. Arkani-Hamed, S. Dimopoulos, and G. Dvali, *Phys. Lett. B* **429**, 263 (1998).
- [2] N. Arkani-Hamed, S. Dimopoulos, and G. Dvali, *Phys. Rev. D* **59**, 086004 (1999).
- [3] E. Witten, *Phys. Lett. B* **105**, 267 (1981).
- [4] L. Randall and R. Sundrum, *Phys. Rev. Lett.* **83**, 3370 (1999).
- [5] H. Davoudiasl, J. L. Hewett, and T. G. Rizzo, *Phys. Rev. D* **63**, 075004 (2001).
- [6] S. Chatrchyan *et al.* (CMS Collaboration), *J. High Energy Phys.* **05** (2011) 85.
- [7] G. Giudice, R. Rattazzi, and J. Wells, *Nucl. Phys.* **B544**, 3 (1999).
- [8] J. L. Hewett, *Phys. Rev. Lett.* **82**, 4765 (1999).
- [9] T. Han, J. D. Lykken, and R.-J. Zhang, *Phys. Rev. D* **59**, 105006 (1999).
- [10] K. Nakamura *et al.* (Particle Data Group), *J. Phys. G* **37**, 075021 (2010).
- [11] K. Cheung and G. Landsberg, *Phys. Rev. D* **62**, 076003 (2000).
- [12] V. M. Abazov *et al.* (D0 Collaboration), *Phys. Rev. Lett.* **103**, 191803 (2009).
- [13] V. M. Abazov *et al.* (D0 Collaboration), *Phys. Rev. Lett.* **102**, 051601 (2009).
- [14] G. Aad *et al.* (ATLAS Collaboration), *Phys. Rev. Lett.* **107**, 272002 (2011).
- [15] S. Chatrchyan *et al.* (CMS Collaboration), *JINST* **3**, S08004 (2008).
- [16] CMS Collaboration, Report No. CMS-PAS-EWK-10-004, 2010 [<https://cdsweb.cern.ch/record/1279145?ln=en>].
- [17] V. Khachatryan *et al.* (CMS Collaboration), *Eur. Phys. J. C* **70**, 1165 (2010).
- [18] V. Khachatryan *et al.* (CMS Collaboration), *Phys. Rev. Lett.* **106**, 082001 (2011).
- [19] CMS Collaboration, Report No. CMS-PAS-EGM-10-006, 2010 [<https://cdsweb.cern.ch/record/1324545?ln=en>].
- [20] T. Gleisberg *et al.*, *J. High Energy Phys.* **02** (2004) 056.
- [21] M. Kumar, P. Mathews, V. Ravindran, and A. Tripathi, *Phys. Lett. B* **672**, 45 (2009).
- [22] M. Kumar, P. Mathews, V. Ravindran, and A. Tripathi, *Nucl. Phys.* **B818**, 28 (2009).
- [23] T. Sjostrand, P. Eden, C. Friberg, L. Lonnblad, G. Miu, S. Mrenna, and E. Norrbin, *Comput. Phys. Commun.* **135**, 238 (2001).
- [24] P. Nadolsky *et al.* (CTEQ Collaboration), *Phys. Rev. D* **78**, 013004 (2008).
- [25] T. Binoth, J. P. Guillet, E. Pilon, and M. Werlen, *Eur. Phys. J. C* **16**, 311 (2000).
- [26] Z. Bern, L. Dixon, and C. Schmidt, *Phys. Rev. D* **66**, 074018 (2002).
- [27] S. Chatrchyan *et al.* (CMS Collaboration), *J. High Energy Phys.* **01** (2012) 133.
- [28] A. L. Read, *J. Phys. G* **28**, 2693 (2002).
- [29] T. Junk, *Nucl. Instrum. Methods Phys. Res., Sect. A* **434**, 435 (1999).

S. Chatrchyan,¹ V. Khachatryan,¹ A. M. Sirunyan,¹ A. Tumasyan,¹ W. Adam,² T. Bergauer,² M. Dragicevic,² J. Erö,² C. Fabjan,² M. Friedl,² R. Frühwirth,² V. M. Ghete,² J. Hammer,^{2,b} M. Hoch,² N. Hörmann,² J. Hrubec,² M. Jeitler,² W. Kiesenhofer,² A. Knapitsch,² M. Krammer,² D. Liko,² I. Mikulec,² M. Pernicka,^{2,a} B. Rabbaran,² H. Rohringer,² R. Schöfbeck,² J. Strauss,² A. Taurok,² F. Teischinger,² C. Trauner,² P. Wagner,² W. Waltenberger,² G. Walzel,² E. Widl,² C.-E. Wulz,² V. Mossolov,³ N. Shumeiko,³ J. Suarez Gonzalez,³ S. Bansal,⁴ L. Benucci,⁴ E. A. De Wolf,⁴ X. Janssen,⁴ S. Luyckx,⁴ T. Maes,⁴ L. Mucibello,⁴ S. Ochesanu,⁴ B. Roland,⁴ R. Rougny,⁴ M. Selvaggi,⁴ H. Van Haevermaet,⁴ P. Van Mechelen,⁴ N. Van Remortel,⁴ A. Van Spilbeeck,⁴ F. Blekman,⁵ S. Blyweert,⁵ J. D'Hondt,⁵ R. Gonzalez Suarez,⁵ A. Kalogeropoulos,⁵ M. Maes,⁵ A. Olbrechts,⁵ W. Van Doninck,⁵ P. Van Mulders,⁵ G. P. Van Onsem,⁵ I. Villella,⁵ O. Charaf,⁶ B. Clerbaux,⁶ G. De Lentdecker,⁶ V. Dero,⁶ A. P. R. Gay,⁶ G. H. Hammad,⁶ T. Hreus,⁶ A. Léonard,⁶ P. E. Marage,⁶ L. Thomas,⁶ C. Vander Velde,⁶ P. Vanlaer,⁶ J. Wickens,⁶ V. Adler,⁷ K. Beernaert,⁷ A. Cimmino,⁷ S. Costantini,⁷ M. Grunewald,⁷ B. Klein,⁷ J. Lellouch,⁷ A. Marinov,⁷ J. McCartin,⁷ D. Ryckbosch,⁷ N. Strobbe,⁷ F. Thyssen,⁷ M. Tytgat,⁷ L. Vanelderen,⁷ P. Verwilligen,⁷ S. Walsh,⁷ N. Zaganidis,⁷ S. Bassegmez,⁸ G. Bruno,⁸ J. Caudron,⁸ L. Ceard,⁸ J. De Favereau De Jeneret,⁸ C. Delaere,⁸ D. Favart,⁸ L. Forthomme,⁸ A. Giammanco,^{8,c} G. Grégoire,⁸ J. Hollar,⁸ V. Lemaître,⁸ J. Liao,⁸ O. Militaru,⁸ C. Nuttens,⁸ D. Pagano,⁸ A. Pin,⁸ K. Piotrkowski,⁸ N. Schul,⁸ N. Beliy,⁹ T. Caebergs,⁹ E. Daubie,⁹ G. A. Alves,¹⁰ D. De Jesus Damiao,¹⁰ M. E. Pol,¹⁰ M. H. G. Souza,¹⁰ W. L. Aldá Júnior,¹¹ W. Carvalho,¹¹ A. Custódio,¹¹ E. M. Da Costa,¹¹ C. De Oliveira Martins,¹¹ S. Fonseca De Souza,¹¹ D. Matos Figueiredo,¹¹ L. Mundim,¹¹ H. Nogima,¹¹ V. Oguri,¹¹ W. L. Prado Da Silva,¹¹ A. Santoro,¹¹ S. M. Silva Do Amaral,¹¹ A. Sznajder,¹¹ T. S. Anjos,^{12,d} C. A. Bernardes,^{12,d} F. A. Dias,^{12,e} T. R. Fernandez Perez Tomei,¹² E. M. Gregores,^{12,d} C. Lagana,¹² F. Marinho,¹² P. G. Mercadante,^{12,d} S. F. Novaes,¹² Sandra S. Padula,¹² N. Dardanov,^{13,b} V. Genchev,^{13,b} P. Iaydjiev,^{13,b} S. Piperov,¹³ M. Rodozov,¹³ S. Stoykova,¹³ G. Sultanov,¹³ V. Tcholakov,¹³ R. Trayanov,¹³ M. Vutova,¹³ A. Dimitrov,¹⁴ R. Hadjiiska,¹⁴ A. Karadzhinova,¹⁴ V. Kozhuharov,¹⁴ L. Litov,¹⁴ B. Pavlov,¹⁴ P. Petkov,¹⁴ J. G. Bian,¹⁵ G. M. Chen,¹⁵ H. S. Chen,¹⁵ C. H. Jiang,¹⁵ D. Liang,¹⁵ S. Liang,¹⁵ X. Meng,¹⁵ J. Tao,¹⁵ J. Wang,¹⁵ J. Wang,¹⁵ X. Wang,¹⁵ Z. Wang,¹⁵ H. Xiao,¹⁵ M. Xu,¹⁵ J. Zang,¹⁵ Z. Zhang,¹⁵ Y. Ban,¹⁶ S. Guo,¹⁶ Y. Guo,¹⁶ W. Li,¹⁶ S. Liu,¹⁶ Y. Mao,¹⁶ S. J. Qian,¹⁶ H. Teng,¹⁶ S. Wang,¹⁶ B. Zhu,¹⁶ W. Zou,¹⁶ A. Cabrera,¹⁷

B. Gomez Moreno,¹⁷ A. A. Ocampo Rios,¹⁷ A. F. Osorio Oliveros,¹⁷ J. C. Sanabria,¹⁷ N. Godinovic,¹⁸ D. Lelas,¹⁸ R. Plestina,^{18,f} D. Polic,¹⁸ I. Puljak,^{18,b} Z. Antunovic,¹⁹ M. Dzelalija,¹⁹ M. Kovac,¹⁹ V. Brigljevic,²⁰ S. Duric,²⁰ K. Kadija,²⁰ J. Luetic,²⁰ S. Morovic,²⁰ A. Attikis,²¹ M. Galanti,²¹ J. Mousa,²¹ C. Nicolaou,²¹ F. Ptochos,²¹ P. A. Razis,²¹ M. Finger,²² M. Finger, Jr.,²² Y. Assran,^{23,g} A. Ellithi Kamel,^{23,h} S. Khalil,^{23,i} M. A. Mahmoud,^{23,j} A. Radi,^{23,k} A. Hektor,²⁴ M. Kadasik,²⁴ M. Müntel,²⁴ M. Raidal,²⁴ L. Rebane,²⁴ A. Tiko,²⁴ V. Azzolini,²⁵ P. Eerola,²⁵ G. Fedi,²⁵ M. Voutilainen,²⁵ S. Czellar,²⁶ J. Härkönen,²⁶ A. Heikkinen,²⁶ V. Karimäki,²⁶ R. Kinnunen,²⁶ M. J. Kortelainen,²⁶ T. Lampén,²⁶ K. Lassila-Perini,²⁶ S. Lehti,²⁶ T. Lindén,²⁶ P. Luukka,²⁶ T. Mäenpää,²⁶ E. Tuominen,²⁶ J. Tuominiemi,²⁶ E. Tuovinen,²⁶ D. Ungaro,²⁶ L. Wendland,²⁶ K. Banzuzi,²⁷ A. Karjalainen,²⁷ A. Korpela,²⁷ T. Tuuva,²⁷ D. Sillou,²⁸ M. Besancon,²⁹ S. Choudhury,²⁹ M. Dejardin,²⁹ D. Denegri,²⁹ B. Fabbro,²⁹ J. L. Faure,²⁹ F. Ferri,²⁹ S. Ganjour,²⁹ A. Givernaud,²⁹ P. Gras,²⁹ G. Hamel de Monchenault,²⁹ P. Jarry,²⁹ E. Locci,²⁹ J. Malcles,²⁹ M. Marionneau,²⁹ L. Millischer,²⁹ J. Rander,²⁹ A. Rosowsky,²⁹ I. Shreyber,²⁹ M. Titov,²⁹ S. Baffioni,³⁰ F. Beaudette,³⁰ L. Benhabib,³⁰ L. Bianchini,³⁰ M. Bluj,^{30,l} C. Broutin,³⁰ P. Busson,³⁰ C. Charlot,³⁰ N. Daci,³⁰ T. Dahms,³⁰ L. Dobrzynski,³⁰ S. Elgammal,³⁰ R. Granier de Cassagnac,³⁰ M. Haguenaer,³⁰ P. Miné,³⁰ C. Mironov,³⁰ C. Ochando,³⁰ P. Paganini,³⁰ D. Sabes,³⁰ R. Salerno,³⁰ Y. Sirois,³⁰ C. Thiebaux,³⁰ C. Veelken,³⁰ A. Zabi,³⁰ J.-L. Agram,^{31,m} J. Andrea,³¹ D. Bloch,³¹ D. Bodin,³¹ J.-M. Brom,³¹ M. Cardaci,³¹ E. C. Chabert,³¹ C. Collard,³¹ E. Conte,^{31,m} F. Drouhin,^{31,m} C. Ferro,³¹ J.-C. Fontaine,^{31,m} D. Gelé,³¹ U. Goerlach,³¹ S. Greder,³¹ P. Juillot,³¹ M. Karim,^{31,m} A.-C. Le Bihan,³¹ P. Van Hove,³¹ F. Fassi,³² D. Mercier,³² C. Baty,³³ S. Beauceron,³³ N. Beaupere,³³ M. Bedjidian,³³ O. Bondu,³³ G. Boudoul,³³ D. Boumediene,³³ H. Brun,³³ J. Chasserat,³³ R. Chierici,^{33,b} D. Contardo,³³ P. Depasse,³³ H. El Mamouni,³³ A. Falkiewicz,³³ J. Fay,³³ S. Gascon,³³ M. Gouzevitch,³³ B. Ille,³³ T. Kurca,³³ T. Le Grand,³³ M. Lethuillier,³³ L. Mirabito,³³ S. Perries,³³ V. Sordini,³³ S. Tosi,³³ Y. Tschudi,³³ P. Verdier,³³ S. Viret,³³ D. Lomidze,³⁴ G. Anagnostou,³⁵ S. Beranek,³⁵ M. Edelhoff,³⁵ L. Feld,³⁵ N. Heracleous,³⁵ O. Hindrichs,³⁵ R. Jussen,³⁵ K. Klein,³⁵ J. Merz,³⁵ A. Ostapchuk,³⁵ A. Perieanu,³⁵ F. Raupach,³⁵ J. Sammet,³⁵ S. Schael,³⁵ D. Sprenger,³⁵ H. Weber,³⁵ B. Wittmer,³⁵ V. Zhukov,^{35,n} M. Ata,³⁶ E. Dietz-Laursonn,³⁶ M. Erdmann,³⁶ T. Hebbeker,³⁶ C. Heidemann,³⁶ A. Hinzmann,³⁶ K. Hoepfner,³⁶ T. Klimkovich,³⁶ D. Klingebiel,³⁶ P. Kreuzer,³⁶ D. Lanske,^{36,a} J. Lingemann,³⁶ C. Magass,³⁶ M. Merschmeyer,³⁶ A. Meyer,³⁶ P. Papacz,³⁶ H. Pieta,³⁶ H. Reithler,³⁶ S. A. Schmitz,³⁶ L. Sonnenschein,³⁶ J. Steggemann,³⁶ D. Teyssier,³⁶ M. Weber,³⁶ M. Bontenackels,³⁷ V. Cherepanov,³⁷ M. Davids,³⁷ G. Flügge,³⁷ H. Geenen,³⁷ W. Haj Ahmad,³⁷ F. Hoehle,³⁷ B. Kargoll,³⁷ T. Kress,³⁷ Y. Kuessel,³⁷ A. Linn,³⁷ A. Nowack,³⁷ L. Perchalla,³⁷ O. Pooth,³⁷ J. Rennefeld,³⁷ P. Sauerland,³⁷ A. Stahl,³⁷ D. Tornier,³⁷ M. H. Zoeller,³⁷ M. Aldaya Martin,³⁸ W. Behrenhoff,³⁸ U. Behrens,³⁸ M. Bergholz,^{38,o} A. Bethani,³⁸ K. Borras,³⁸ A. Cakir,³⁸ A. Campbell,³⁸ E. Castro,³⁸ D. Dammann,³⁸ G. Eckerlin,³⁸ D. Eckstein,³⁸ A. Flossdorf,³⁸ G. Flucke,³⁸ A. Geiser,³⁸ J. Hauk,³⁸ H. Jung,^{38,b} M. Kasemann,³⁸ P. Katsas,³⁸ C. Kleinwort,³⁸ H. Kluge,³⁸ A. Knutsson,³⁸ M. Krämer,³⁸ D. Krücker,³⁸ E. Kuznetsova,³⁸ W. Lange,³⁸ W. Lohmann,^{38,o} B. Lutz,³⁸ R. Mankel,³⁸ I. Marfin,³⁸ M. Marienfeld,³⁸ I.-A. Melzer-Pellmann,³⁸ A. B. Meyer,³⁸ J. Mnich,³⁸ A. Mussgiller,³⁸ S. Naumann-Emme,³⁸ J. Olzem,³⁸ A. Petrukhin,³⁸ D. Pitzl,³⁸ A. Raspereza,³⁸ M. Rosin,³⁸ J. Salfeld-Nebgen,³⁸ R. Schmidt,^{38,o} T. Schoerner-Sadenius,³⁸ N. Sen,³⁸ A. Spiridonov,³⁸ M. Stein,³⁸ J. Tomaszewska,³⁸ R. Walsh,³⁸ C. Wissing,³⁸ C. Autermann,³⁹ V. Blobel,³⁹ S. Bobrovskiy,³⁹ J. Draeger,³⁹ H. Enderle,³⁹ U. Gebbert,³⁹ M. Görner,³⁹ T. Hermanns,³⁹ K. Kaschube,³⁹ G. Kaussen,³⁹ H. Kirschenmann,³⁹ R. Klanner,³⁹ J. Lange,³⁹ B. Mura,³⁹ F. Nowak,³⁹ N. Pietsch,³⁹ C. Sander,³⁹ H. Schettler,³⁹ P. Schleper,³⁹ E. Schlieckau,³⁹ M. Schröder,³⁹ T. Schum,³⁹ H. Stadie,³⁹ G. Steinbrück,³⁹ J. Thomsen,³⁹ C. Barth,⁴⁰ J. Berger,⁴⁰ T. Chwalek,⁴⁰ W. De Boer,⁴⁰ A. Dierlamm,⁴⁰ G. Dirkes,⁴⁰ M. Feindt,⁴⁰ J. Gruschke,⁴⁰ M. Guthoff,^{40,b} C. Hackstein,⁴⁰ F. Hartmann,⁴⁰ M. Heinrich,⁴⁰ H. Held,⁴⁰ K. H. Hoffmann,⁴⁰ S. Honc,⁴⁰ I. Katkov,^{40,n} J. R. Komaragiri,⁴⁰ T. Kuhr,⁴⁰ D. Martschei,⁴⁰ S. Mueller,⁴⁰ Th. Müller,⁴⁰ M. Niegel,⁴⁰ O. Oberst,⁴⁰ A. Oehler,⁴⁰ J. Ott,⁴⁰ T. Peiffer,⁴⁰ G. Quast,⁴⁰ K. Rabbertz,⁴⁰ F. Ratnikov,⁴⁰ N. Ratnikova,⁴⁰ M. Renz,⁴⁰ S. Röcker,⁴⁰ C. Saout,⁴⁰ A. Scheurer,⁴⁰ P. Schieferdecker,⁴⁰ F.-P. Schilling,⁴⁰ M. Schmanau,⁴⁰ G. Schott,⁴⁰ H. J. Simonis,⁴⁰ F. M. Stober,⁴⁰ D. Troendle,⁴⁰ J. Wagner-Kuhr,⁴⁰ T. Weiler,⁴⁰ M. Zeise,⁴⁰ E. B. Ziebarth,⁴⁰ G. Daskalakis,⁴¹ T. Gerasis,⁴¹ S. Kesisoglou,⁴¹ A. Kyriakis,⁴¹ D. Loukas,⁴¹ I. Manolagos,⁴¹ A. Markou,⁴¹ C. Markou,⁴¹ C. Mavrommatis,⁴¹ E. Ntomari,⁴¹ E. Petrakou,⁴¹ L. Gouskos,⁴² T. J. Mertzimekis,⁴² A. Panagiotou,⁴² N. Saoulidou,⁴² E. Stiliaris,⁴² I. Evangelou,⁴³ C. Foudas,^{43,b} P. Kokkas,⁴³ N. Manthos,⁴³ I. Papadopoulos,⁴³ V. Patras,⁴³ F. A. Triantis,⁴³ A. Aranyi,⁴⁴ G. Bencze,⁴⁴ L. Boldizsar,⁴⁴ C. Hajdu,^{44,b} P. Hidas,⁴⁴ D. Horvath,^{44,p} A. Kapusi,⁴⁴ K. Krajczar,^{44,q} F. Sikler,^{44,b} G. I. Veres,^{44,q} G. Vesztegombi,^{44,q} N. Beni,⁴⁵ J. Molnar,⁴⁵ J. Palinkas,⁴⁵ Z. Szillasi,⁴⁵ V. Veszpremi,⁴⁵ J. Karancsi,⁴⁶ P. Raics,⁴⁶ Z. L. Trocsanyi,⁴⁶ B. Ujvari,⁴⁶ S. B. Beri,⁴⁷ V. Bhatnagar,⁴⁷ N. Dhingra,⁴⁷ R. Gupta,⁴⁷ M. Jindal,⁴⁷ M. Kaur,⁴⁷ J. M. Kohli,⁴⁷ M. Z. Mehta,⁴⁷ N. Nishu,⁴⁷

L. K. Saini,⁴⁷ A. Sharma,⁴⁷ A. P. Singh,⁴⁷ J. Singh,⁴⁷ S. P. Singh,⁴⁷ S. A. Ahuja,⁴⁸ B. C. Choudhary,⁴⁸ A. Kumar,⁴⁸
A. Kumar,⁴⁸ S. Malhotra,⁴⁸ M. Naimuddin,⁴⁸ K. Ranjan,⁴⁸ R. K. Shivpuri,⁴⁸ S. Banerjee,⁴⁹ S. Bhattacharya,⁴⁹
S. Dutta,⁴⁹ B. Gomber,⁴⁹ S. Jain,⁴⁹ S. Jain,⁴⁹ R. Khurana,⁴⁹ S. Sarkar,⁴⁹ R. K. Choudhury,⁵⁰ D. Dutta,⁵⁰ S. Kailas,⁵⁰
V. Kumar,⁵⁰ A. K. Mohanty,^{50,b} L. M. Pant,⁵⁰ P. Shukla,⁵⁰ T. Aziz,⁵¹ M. Guchait,^{51,r} A. Gurtu,^{51,s} M. Maity,^{51,s}
D. Majumder,⁵¹ G. Majumder,⁵¹ K. Mazumdar,⁵¹ G. B. Mohanty,⁵¹ B. Parida,⁵¹ A. Saha,⁵¹ K. Sudhakar,⁵¹
N. Wickramage,⁵¹ S. Banerjee,⁵² S. Dugad,⁵² N. K. Mondal,⁵² H. Arfaei,⁵³ H. Bakhshiansohi,^{53,t} S. M. Etesami,^{53,u}
A. Fahim,^{53,t} M. Hashemi,⁵³ H. Hesari,⁵³ A. Jafari,^{53,t} M. Khakzad,⁵³ A. Mohammadi,^{53,v}
M. Mohammadi Najafabadi,⁵³ S. Paktinat Mehdiabadi,⁵³ B. Safarzadeh,^{53,w} M. Zeinali,^{53,u} M. Abbrescia,^{54a,54b}
L. Barbone,^{54a,54b} C. Calabria,^{54a,54b} A. Colaleo,^{54a} D. Creanza,^{54a,54c} N. De Filippis,^{54a,54c,b} M. De Palma,^{54a,54b}
L. Fiore,^{54a} G. Iaselli,^{54a,54c} L. Lusito,^{54a,54b} G. Maggi,^{54a,54c} M. Maggi,^{54a} N. Manna,^{54a,54b} B. Marangelli,^{54a,54b}
S. My,^{54a,54c} S. Nuzzo,^{54a,54b} N. Pacifico,^{54a,54b} A. Pompili,^{54a,54b} G. Pugliese,^{54a,54c} F. Romano,^{54a,54c}
G. Selvaggi,^{54a,54b} L. Silvestris,^{54a} S. Tuppiti,^{54a,54b} G. Zito,^{54a} G. Abbiendi,^{55a} A. C. Benvenuti,^{55a} D. Bonacorsi,^{55a}
S. Braibant-Giacomelli,^{55a,55b} L. Brigliadori,^{55a} P. Capiluppi,^{55a,55b} A. Castro,^{55a,55b} F. R. Cavallo,^{55a}
M. Cuffiani,^{55a,55b} G. M. Dallavalle,^{55a} F. Fabbri,^{55a} A. Fanfani,^{55a,55b} D. Fasanella,^{55a,b} P. Giacomelli,^{55a}
C. Grandi,^{55a} S. Marcellini,^{55a} G. Masetti,^{55a} M. Meneghelli,^{55a,55b} A. Montanari,^{55a} F. L. Navarria,^{55a,55b}
F. Odorici,^{55a} A. Perrotta,^{55a} F. Primavera,^{55a} A. M. Rossi,^{55a,55b} T. Rovelli,^{55a,55b} G. Siroli,^{55a,55b}
R. Travaglini,^{55a,55b} S. Albergo,^{56a,56b} G. Cappello,^{56a,56b} M. Chiorboli,^{56a,56b} S. Costa,^{56a,56b} R. Potenza,^{56a,56b}
A. Tricomi,^{56a,56b} C. Tuve,^{56a,56b} G. Barbagli,^{57a} V. Ciulli,^{57a,57b} C. Civinini,^{57a} R. D'Alessandro,^{57a,57b}
E. Focardi,^{57a,57b} S. Frosali,^{57a,57b} E. Gallo,^{57a} S. Gonzi,^{57a,57b} M. Meschini,^{57a} S. Paoletti,^{57a} G. Sguazzoni,^{57a}
A. Tropiano,^{57a,b} L. Benussi,⁵⁸ S. Bianco,⁵⁸ S. Colafranceschi,^{58,x} F. Fabbri,⁵⁸ D. Piccolo,⁵⁸ P. Fabbricatore,⁵⁹
R. Musenich,⁵⁹ A. Benaglia,^{60a,60b} F. De Guio,^{60a,60b} L. Di Matteo,^{60a,60b} S. Gennai,^{60a,b} A. Ghezzi,^{60a,60b}
S. Malvezzi,^{60a} A. Martelli,^{60a,60b} A. Massironi,^{60a,60b} D. Menasce,^{60a} L. Moroni,^{60a} M. Paganoni,^{60a,60b}
D. Pedrini,^{60a} S. Ragazzi,^{60a,60b} N. Redaelli,^{60a} S. Sala,^{60a} T. Tabarelli de Fatis,^{60a,60b} S. Buontempo,^{61a}
C. A. Carrillo Montoya,^{61a,b} N. Cavallo,^{61a,y} A. De Cosa,^{61a,61b} O. Dogangun,^{61a,61b} F. Fabozzi,^{61a,y}
A. O. M. Iorio,^{61a,b} L. Lista,^{61a} M. Merola,^{61a,61b} P. Paolucci,^{61a} P. Azzi,^{62a} N. Bacchetta,^{62a,b} P. Bellan,^{62a,62b}
D. Bisello,^{62a,62b} A. Branca,^{62a} R. Carlin,^{62a,62b} P. Checchia,^{62a} T. Dorigo,^{62a} U. Dosselli,^{62a,62c} F. Fanzago,^{62a}
F. Gasparini,^{62a,62b} U. Gasparini,^{62a} A. Gozzelino,^{62a} S. Lacaprara,^{62a,z} I. Lazzizzera,^{62a,62c} M. Margoni,^{62a,62b}
M. Mazzucato,^{62a} A. T. Meneguzzo,^{62a,62b} M. Nespolo,^{62a,b} L. Perrozzi,^{62a} N. Pozzobon,^{62a,62b} P. Ronchese,^{62a,62b}
F. Simonetto,^{62a,62b} E. Torassa,^{62a} M. Tosi,^{62a,62b} S. Vanini,^{62a,62b} P. Zotto,^{62a,62b} G. Zumerle,^{62a,62b} P. Baesso,^{63a,63b}
U. Berzano,^{63a} S. P. Ratti,^{63a,63b} C. Riccardi,^{63a,63b} P. Torre,^{63a,63b} P. Vitulo,^{63a,63b} C. Viviani,^{63a,63b} M. Biasini,^{64a,64b}
G. M. Bilei,^{64a} B. Caponeri,^{64a,64b} L. Fanò,^{64a,64b} P. Lariccia,^{64a,64b} A. Lucaroni,^{64a,64b} G. Mantovani,^{64a,64b}
M. Menichelli,^{64a} A. Nappi,^{64a,64b} F. Romeo,^{64a,64b} A. Santocchia,^{64a,64b} S. Taroni,^{64a,64b} M. Valdata,^{64a,64b}
P. Azzurri,^{65a,65c} G. Bagliesi,^{65a} T. Boccali,^{65a} G. Broccolo,^{65a,65c} R. Castaldi,^{65a} R. T. D'Agnolo,^{65a,65c}
R. Dell'Orso,^{65a} F. Fiori,^{65a,65b} L. Foà,^{65a,65c} A. Giassi,^{65a} A. Kraan,^{65a} F. Ligabue,^{65a,65c} T. Lomtadze,^{65a}
L. Martini,^{65a,aa} A. Messineo,^{65a,65b} F. Palla,^{65a} F. Palmonari,^{65a} A. Rizzi,^{65a} G. Segneri,^{65a} A. T. Serban,^{65a}
P. Spagnolo,^{65a} R. Tenchini,^{65a} G. Tonelli,^{65a,65b} A. Venturi,^{65a,b} P. G. Verdini,^{65a} L. Barone,^{66a,66b} F. Cavallari,^{66a}
D. Del Re,^{66a,66b} M. Diemoz,^{66a} D. Franci,^{66a,66b} M. Grassi,^{66a,b} E. Longo,^{66a,66b} P. Meridiani,^{66a} S. Nourbakhsh,^{66a}
G. Organtini,^{66a,66b} F. Pandolfi,^{66a,66b} R. Paramatti,^{66a} S. Rahatlou,^{66a,66b} M. Sigamani,^{66a} N. Amapane,^{67a,67b}
R. Arcidiacono,^{67a,67c} S. Argiro,^{67a,67b} M. Arneodo,^{67a,67c} C. Biino,^{67a} C. Botta,^{67a,67b} N. Cartiglia,^{67a}
R. Castello,^{67a,67b} M. Costa,^{67a,67b} N. Demaria,^{67a} A. Graziano,^{67a,67b} C. Mariotti,^{67a,b} S. Maselli,^{67a}
E. Migliore,^{67a,67b} V. Monaco,^{67a,67b} M. Musich,^{67a} M. M. Obertino,^{67a,67c} N. Pastrone,^{67a} M. Pelliccioni,^{67a}
A. Potenza,^{67a,67b} A. Romero,^{67a,67b} M. Ruspa,^{67a,67c} R. Sacchi,^{67a,67b} V. Sola,^{67a,67b} A. Solano,^{67a,67b} A. Staiano,^{67a}
A. Vilela Pereira,^{67a} S. Belforte,^{68a} F. Cossutti,^{68a} G. Della Ricca,^{68a,68b} B. Gobbo,^{68a} M. Marone,^{68a,68b}
D. Montanino,^{68a,68b} A. Penzo,^{68a} S. G. Heo,⁶⁹ S. K. Nam,⁶⁹ S. Chang,⁷⁰ J. Chung,⁷⁰ D. H. Kim,⁷⁰ G. N. Kim,⁷⁰
J. E. Kim,⁷⁰ D. J. Kong,⁷⁰ H. Park,⁷⁰ S. R. Ro,⁷⁰ D. C. Son,⁷⁰ T. Son,⁷⁰ J. Y. Kim,⁷¹ Zero J. Kim,⁷¹ S. Song,⁷¹
H. Y. Jo,⁷² S. Choi,⁷³ D. Gyun,⁷³ B. Hong,⁷³ M. Jo,⁷³ H. Kim,⁷³ T. J. Kim,⁷³ K. S. Lee,⁷³ D. H. Moon,⁷³ S. K. Park,⁷³
E. Seo,⁷³ K. S. Sim,⁷³ M. Choi,⁷⁴ S. Kang,⁷⁴ H. Kim,⁷⁴ J. H. Kim,⁷⁴ C. Park,⁷⁴ I. C. Park,⁷⁴ S. Park,⁷⁴ G. Ryu,⁷⁴
Y. Cho,⁷⁵ Y. Choi,⁷⁵ Y. K. Choi,⁷⁵ J. Goh,⁷⁵ M. S. Kim,⁷⁵ B. Lee,⁷⁵ J. Lee,⁷⁵ S. Lee,⁷⁵ H. Seo,⁷⁵ I. Yu,⁷⁵
M. J. Bilinskas,⁷⁶ I. Grigelionis,⁷⁶ M. Janulis,⁷⁶ D. Martisiute,⁷⁶ P. Petrov,⁷⁶ M. Polujanskas,⁷⁶ T. Sabonis,⁷⁶
H. Castilla-Valdez,⁷⁷ E. De La Cruz-Burelo,⁷⁷ I. Heredia-de La Cruz,⁷⁷ R. Lopez-Fernandez,⁷⁷ R. Magaña Villalba,⁷⁷
J. Martínez-Ortega,⁷⁷ A. Sánchez-Hernández,⁷⁷ L. M. Villasenor-Cendejas,⁷⁷ S. Carrillo Moreno,⁷⁸

F. Vazquez Valencia,⁷⁸ H. A. Salazar Ibarguen,⁷⁹ E. Casimiro Linares,⁸⁰ A. Morelos Pineda,⁸⁰ M. A. Reyes-Santos,⁸⁰ D. Krofcheck,⁸¹ A. J. Bell,⁸² P. H. Butler,⁸² R. Doesburg,⁸² S. Reucroft,⁸² H. Silverwood,⁸² M. Ahmad,⁸³ M. I. Asghar,⁸³ H. R. Hoorani,⁸³ S. Khalid,⁸³ W. A. Khan,⁸³ T. Khurshid,⁸³ S. Qazi,⁸³ M. A. Shah,⁸³ M. Shoaib,⁸³ G. Brona,⁸⁴ M. Cwiok,⁸⁴ W. Dominik,⁸⁴ K. Doroba,⁸⁴ A. Kalinowski,⁸⁴ M. Konecki,⁸⁴ J. Krolkowski,⁸⁴ H. Bialkowska,⁸⁵ B. Boimska,⁸⁵ T. Frueboes,⁸⁵ R. Gokiel,⁸⁵ M. Górski,⁸⁵ M. Kazana,⁸⁵ K. Nawrocki,⁸⁵ K. Romanowska-Rybinska,⁸⁵ M. Szeleper,⁸⁵ G. Wrochna,⁸⁵ P. Zalewski,⁸⁵ N. Almeida,⁸⁶ P. Bargassa,⁸⁶ A. David,⁸⁶ P. Faccioli,⁸⁶ P. G. Ferreira Parracho,⁸⁶ M. Gallinaro,⁸⁶ P. Musella,⁸⁶ A. Nayak,⁸⁶ J. Pela,^{86,b} P. Q. Ribeiro,⁸⁶ J. Seixas,⁸⁶ J. Varela,⁸⁶ S. Afanasiev,⁸⁷ I. Belotelov,⁸⁷ P. Bunin,⁸⁷ M. Gavrilenko,⁸⁷ I. Golutvin,⁸⁷ I. Gorbunov,⁸⁷ A. Kamenev,⁸⁷ V. Karjavin,⁸⁷ G. Kozlov,⁸⁷ A. Lanev,⁸⁷ P. Moisenz,⁸⁷ V. Palichik,⁸⁷ V. Perelygin,⁸⁷ S. Shmatov,⁸⁷ V. Smirnov,⁸⁷ A. Volodko,⁸⁷ A. Zarubin,⁸⁷ S. Evstyukhin,⁸⁸ V. Golovtsov,⁸⁸ Y. Ivanov,⁸⁸ V. Kim,⁸⁸ P. Levchenko,⁸⁸ V. Murzin,⁸⁸ V. Oreshkin,⁸⁸ I. Smirnov,⁸⁸ V. Sulimov,⁸⁸ L. Uvarov,⁸⁸ S. Vavilov,⁸⁸ A. Vorobyev,⁸⁸ An. Vorobyev,⁸⁸ Yu. Andreev,⁸⁹ A. Dermenev,⁸⁹ S. Gninenko,⁸⁹ N. Golubev,⁸⁹ M. Kirsanov,⁸⁹ N. Krasnikov,⁸⁹ V. Matveev,⁸⁹ A. Pashenkov,⁸⁹ A. Toropin,⁸⁹ S. Troitsky,⁸⁹ V. Epshteyn,⁹⁰ M. Erofeeva,⁹⁰ V. Gavrilo, ⁹⁰ M. Kossov,^{90,b} A. Krokhotin,⁹⁰ N. Lychkovskaya,⁹⁰ V. Popov,⁹⁰ G. Safronov,⁹⁰ S. Semenov,⁹⁰ V. Stolin,⁹⁰ E. Vlasov,⁹⁰ A. Zhokin,⁹⁰ A. Belyaev,⁹¹ E. Boos,⁹¹ M. Dubinin,^{91,e} L. Dudko,⁹¹ A. Ershov,⁹¹ A. Gribushin,⁹¹ O. Kodolova,⁹¹ I. Lokhtin,⁹¹ A. Markina,⁹¹ S. Obraztsov,⁹¹ M. Perfilov,⁹¹ S. Petrushanko,⁹¹ L. Sarycheva,⁹¹ V. Savrin,⁹¹ A. Snigirev,⁹¹ V. Andreev,⁹² M. Azarkin,⁹² I. Dremin,⁹² M. Kirakosyan,⁹² A. Leonidov,⁹² G. Mesyats,⁹² S. V. Rusakov,⁹² A. Vinogradov,⁹² I. Azhgirey,⁹³ I. Bayshev,⁹³ S. Bitioukov,⁹³ V. Grishin,^{93,b} V. Kachanov,⁹³ D. Konstantinov,⁹³ A. Korablev,⁹³ V. Krychkin,⁹³ V. Petrov,⁹³ R. Rytun,⁹³ A. Sobol,⁹³ L. Tourtchanovitch,⁹³ S. Troshin,⁹³ N. Tyurin,⁹³ A. Uzunian,⁹³ A. Volkov,⁹³ P. Adzic,^{94,bb} M. Djordjevic,⁹⁴ M. Ekmedzic,⁹⁴ D. Krpic,^{94,bb} J. Milosevic,⁹⁴ M. Aguilar-Benitez,⁹⁵ J. Alcaraz Maestre,⁹⁵ P. Arce,⁹⁵ C. Battilana,⁹⁵ E. Calvo,⁹⁵ M. Cerrada,⁹⁵ M. Chamizo Llatas,⁹⁵ N. Colino,⁹⁵ B. De La Cruz,⁹⁵ A. Delgado Peris,⁹⁵ C. Diez Pardos,⁹⁵ D. Domínguez Vázquez,⁹⁵ C. Fernandez Bedoya,⁹⁵ J. P. Fernández Ramos,⁹⁵ A. Ferrando,⁹⁵ J. Flix,⁹⁵ M. C. Fouz,⁹⁵ P. Garcia-Abia,⁹⁵ O. Gonzalez Lopez,⁹⁵ S. Goy Lopez,⁹⁵ J. M. Hernandez,⁹⁵ M. I. Josa,⁹⁵ G. Merino,⁹⁵ J. Puerta Pelayo,⁹⁵ I. Redondo,⁹⁵ L. Romero,⁹⁵ J. Santaolalla,⁹⁵ M. S. Soares,⁹⁵ C. Willmott,⁹⁵ C. Albajar,⁹⁶ G. Codispoti,⁹⁶ J. F. de Trocóniz,⁹⁶ J. Cuevas,⁹⁷ J. Fernandez Menendez,⁹⁷ S. Folgueras,⁹⁷ I. Gonzalez Caballero,⁹⁷ L. Lloret Iglesias,⁹⁷ J. M. Vizán Garcia,⁹⁷ J. A. Brochero Cifuentes,⁹⁸ I. J. Cabrillo,⁹⁸ A. Calderon,⁹⁸ S. H. Chuang,⁹⁸ J. Duarte Campderros,⁹⁸ M. Felcini,^{98,cc} M. Fernandez,⁹⁸ G. Gomez,⁹⁸ J. Gonzalez Sanchez,⁹⁸ C. Jorda,⁹⁸ P. Lobelle Pardo,⁹⁸ A. Lopez Virto,⁹⁸ J. Marco,⁹⁸ R. Marco,⁹⁸ C. Martinez Rivero,⁹⁸ F. Matorras,⁹⁸ F. J. Munoz Sanchez,⁹⁸ J. Piedra Gomez,^{98,dd} T. Rodrigo,⁹⁸ A. Y. Rodríguez-Marrero,⁹⁸ A. Ruiz-Jimeno,⁹⁸ L. Scodellaro,⁹⁸ M. Sobron Sanudo,⁹⁸ I. Vila,⁹⁸ R. Vilar Cortabitarte,⁹⁸ D. Abbaneo,⁹⁹ E. Auffray,⁹⁹ G. Auzinger,⁹⁹ P. Baillon,⁹⁹ A. H. Ball,⁹⁹ D. Barney,⁹⁹ C. Bernet,^{99,f} W. Bialas,⁹⁹ P. Bloch,⁹⁹ A. Bocci,⁹⁹ H. Breuker,⁹⁹ K. Bunkowski,⁹⁹ T. Camporesi,⁹⁹ G. Cerminara,⁹⁹ T. Christiansen,⁹⁹ J. A. Coarasa Perez,⁹⁹ B. Curé,⁹⁹ D. D'Enterria,⁹⁹ A. De Roeck,⁹⁹ S. Di Guida,⁹⁹ M. Dobson,⁹⁹ N. Dupont-Sagorin,⁹⁹ A. Elliott-Peisert,⁹⁹ B. Frisch,⁹⁹ W. Funk,⁹⁹ A. Gaddi,⁹⁹ G. Georgiou,⁹⁹ H. Gerwig,⁹⁹ M. Giffels,⁹⁹ D. Gigi,⁹⁹ K. Gill,⁹⁹ D. Giordano,⁹⁹ M. Giunta,⁹⁹ F. Glege,⁹⁹ R. Gomez-Reino Garrido,⁹⁹ P. Govoni,⁹⁹ S. Gowdy,⁹⁹ R. Guida,⁹⁹ L. Guiducci,⁹⁹ S. Gundacker,⁹⁹ M. Hansen,⁹⁹ C. Hartl,⁹⁹ J. Harvey,⁹⁹ J. Hegeman,⁹⁹ B. Hegner,⁹⁹ H. F. Hoffmann,⁹⁹ V. Innocente,⁹⁹ P. Janot,⁹⁹ K. Kaadze,⁹⁹ E. Karavakis,⁹⁹ P. Lecoq,⁹⁹ P. Lenzi,⁹⁹ C. Lourenço,⁹⁹ T. Mäki,⁹⁹ M. Malberti,⁹⁹ L. Malgeri,⁹⁹ M. Mannelli,⁹⁹ L. Masetti,⁹⁹ G. Mavromanolakis,⁹⁹ F. Meijers,⁹⁹ S. Mersi,⁹⁹ E. Meschi,⁹⁹ R. Moser,⁹⁹ M. U. Mozer,⁹⁹ M. Mulders,⁹⁹ E. Nesvold,⁹⁹ M. Nguyen,⁹⁹ T. Orimoto,⁹⁹ L. Orsini,⁹⁹ E. Palencia Cortezon,⁹⁹ E. Perez,⁹⁹ A. Petrilli,⁹⁹ A. Pfeiffer,⁹⁹ M. Pierini,⁹⁹ M. Pimiä,⁹⁹ D. Piparo,⁹⁹ G. Polese,⁹⁹ L. Quertenmont,⁹⁹ A. Racz,⁹⁹ W. Reece,⁹⁹ J. Rodrigues Antunes,⁹⁹ G. Rolandi,^{99,ee} T. Rommerskirchen,⁹⁹ C. Rovelli,^{99,ff} M. Rovere,⁹⁹ H. Sakulin,⁹⁹ F. Santanastasio,⁹⁹ C. Schäfer,⁹⁹ C. Schwick,⁹⁹ I. Segoni,⁹⁹ A. Sharma,⁹⁹ P. Siegrist,⁹⁹ P. Silva,⁹⁹ M. Simon,⁹⁹ P. Sphicas,^{99,gg} D. Spiga,⁹⁹ M. Spiropulu,^{99,e} M. Stoye,⁹⁹ A. Tsirou,⁹⁹ P. Vichoudis,⁹⁹ H. K. Wöhri,⁹⁹ S. D. Worm,^{99,hh} W. D. Zeuner,⁹⁹ W. Bertl,¹⁰⁰ K. Deiters,¹⁰⁰ W. Erdmann,¹⁰⁰ K. Gabathuler,¹⁰⁰ R. Horisberger,¹⁰⁰ Q. Ingram,¹⁰⁰ H. C. Kaestli,¹⁰⁰ S. König,¹⁰⁰ D. Kotlinski,¹⁰⁰ U. Langenegger,¹⁰⁰ F. Meier,¹⁰⁰ D. Renker,¹⁰⁰ T. Rohe,¹⁰⁰ J. Sibille,^{100,ii} L. Bäni,¹⁰¹ P. Bortignon,¹⁰¹ B. Casal,¹⁰¹ N. Chanon,¹⁰¹ Z. Chen,¹⁰¹ S. Cittolin,¹⁰¹ A. Deisher,¹⁰¹ G. Dissertori,¹⁰¹ M. Dittmar,¹⁰¹ J. Eugster,¹⁰¹ K. Freudenreich,¹⁰¹ C. Grab,¹⁰¹ P. Lecomte,¹⁰¹ W. Lustermann,¹⁰¹ P. Martinez Ruiz del Arbol,¹⁰¹ P. Milenovic,^{101,jj} N. Mohr,¹⁰¹ F. Moortgat,¹⁰¹ C. Nägeli,^{101,kk} P. Nef,¹⁰¹ F. Nessi-Tedaldi,¹⁰¹ L. Pape,¹⁰¹ F. Pauss,¹⁰¹ M. Peruzzi,¹⁰¹ F. J. Ronga,¹⁰¹ M. Rossini,¹⁰¹ L. Sala,¹⁰¹ A. K. Sanchez,¹⁰¹ M.-C. Sawley,¹⁰¹ A. Starodumov,^{101,ll} B. Stieger,¹⁰¹ M. Takahashi,¹⁰¹ L. Tauscher,^{101,a} A. Thea,¹⁰¹

K. Theofilatos,¹⁰¹ D. Treille,¹⁰¹ C. Urscheler,¹⁰¹ R. Wallny,¹⁰¹ H. A. Weber,¹⁰¹ L. Wehrli,¹⁰¹ J. Weng,¹⁰¹ E. Aguilo,¹⁰² C. Amsler,¹⁰² V. Chiochia,¹⁰² S. De Visscher,¹⁰² C. Favaro,¹⁰² M. Ivova Rikova,¹⁰² B. Millan Mejias,¹⁰² P. Otiougova,¹⁰² P. Robmann,¹⁰² A. Schmidt,¹⁰² H. Snoek,¹⁰² M. Verzetti,¹⁰² Y. H. Chang,¹⁰³ K. H. Chen,¹⁰³ C. M. Kuo,¹⁰³ S. W. Li,¹⁰³ W. Lin,¹⁰³ Z. K. Liu,¹⁰³ Y. J. Lu,¹⁰³ D. Mekterovic,¹⁰³ R. Volpe,¹⁰³ S. S. Yu,¹⁰³ P. Bartalini,¹⁰⁴ P. Chang,¹⁰⁴ Y. H. Chang,¹⁰⁴ Y. W. Chang,¹⁰⁴ Y. Chao,¹⁰⁴ K. F. Chen,¹⁰⁴ C. Dietz,¹⁰⁴ U. Grundler,¹⁰⁴ W.-S. Hou,¹⁰⁴ Y. Hsiung,¹⁰⁴ K. Y. Kao,¹⁰⁴ Y. J. Lei,¹⁰⁴ R.-S. Lu,¹⁰⁴ J. G. Shiu,¹⁰⁴ Y. M. Tzeng,¹⁰⁴ X. Wan,¹⁰⁴ M. Wang,¹⁰⁴ A. Adiguzel,¹⁰⁵ M. N. Bakirci,^{105,mm} S. Cerci,^{105,nn} C. Dozen,¹⁰⁵ I. Dumanoglu,¹⁰⁵ E. Eskut,¹⁰⁵ S. Girgis,¹⁰⁵ G. Gokbulut,¹⁰⁵ I. Hos,¹⁰⁵ E. E. Kangal,¹⁰⁵ G. Karapinar,¹⁰⁵ A. Kayis Topaksu,¹⁰⁵ G. Onengut,¹⁰⁵ K. Ozdemir,¹⁰⁵ S. Ozturk,^{105,oo} A. Polatoz,¹⁰⁵ K. Sogut,^{105,pp} D. Sunar Cerci,^{105,nn} B. Tali,^{105,nn} H. Topakli,^{105,mm} D. Uzun,¹⁰⁵ L. N. Vergili,¹⁰⁵ M. Vergili,¹⁰⁵ I. V. Akin,¹⁰⁶ T. Aliev,¹⁰⁶ B. Bilin,¹⁰⁶ S. Bilmis,¹⁰⁶ M. Deniz,¹⁰⁶ H. Gamsizkan,¹⁰⁶ A. M. Guler,¹⁰⁶ K. Ocalan,¹⁰⁶ A. Ozpineci,¹⁰⁶ M. Serin,¹⁰⁶ R. Sever,¹⁰⁶ U. E. Surat,¹⁰⁶ M. Yalvac,¹⁰⁶ E. Yildirim,¹⁰⁶ M. Zeyrek,¹⁰⁶ M. Deliomeroğlu,¹⁰⁷ E. Gülmez,¹⁰⁷ B. Isildak,¹⁰⁷ M. Kaya,^{107,qq} O. Kaya,^{107,qq} M. Özbek,¹⁰⁷ S. Ozkorucuklu,^{107,rr} N. Sonmez,^{107,ss} L. Levchuk,¹⁰⁸ F. Bostock,¹⁰⁹ J. J. Brooke,¹⁰⁹ E. Clement,¹⁰⁹ D. Cussans,¹⁰⁹ R. Frazier,¹⁰⁹ J. Goldstein,¹⁰⁹ M. Grimes,¹⁰⁹ G. P. Heath,¹⁰⁹ H. F. Heath,¹⁰⁹ L. Kreczko,¹⁰⁹ S. Metson,¹⁰⁹ D. M. Newbold,^{109,hh} K. Nirunpong,¹⁰⁹ A. Poll,¹⁰⁹ S. Senkin,¹⁰⁹ V. J. Smith,¹⁰⁹ T. Williams,¹⁰⁹ L. Basso,^{110,tt} K. W. Bell,¹¹⁰ A. Belyaev,^{110,tt} C. Brew,¹¹⁰ R. M. Brown,¹¹⁰ B. Camanzi,¹¹⁰ D. J. A. Cockerill,¹¹⁰ J. A. Coughlan,¹¹⁰ K. Harder,¹¹⁰ S. Harper,¹¹⁰ J. Jackson,¹¹⁰ B. W. Kennedy,¹¹⁰ E. Olaiya,¹¹⁰ D. Petyt,¹¹⁰ B. C. Radburn-Smith,¹¹⁰ C. H. Shepherd-Themistocleous,¹¹⁰ I. R. Tomalin,¹¹⁰ W. J. Womersley,¹¹⁰ R. Bainbridge,¹¹¹ G. Ball,¹¹¹ R. Beuselinck,¹¹¹ O. Buchmuller,¹¹¹ D. Colling,¹¹¹ N. Cripps,¹¹¹ M. Cutajar,¹¹¹ P. Dauncey,¹¹¹ G. Davies,¹¹¹ M. Della Negra,¹¹¹ W. Ferguson,¹¹¹ J. Fulcher,¹¹¹ D. Futyan,¹¹¹ A. Gilbert,¹¹¹ A. Guneratne Bryer,¹¹¹ G. Hall,¹¹¹ Z. Hatherell,¹¹¹ J. Hays,¹¹¹ G. Iles,¹¹¹ M. Jarvis,¹¹¹ G. Karapostoli,¹¹¹ L. Lyons,¹¹¹ A.-M. Magnan,¹¹¹ J. Marrouche,¹¹¹ B. Mathias,¹¹¹ R. Nandi,¹¹¹ J. Nash,¹¹¹ A. Nikitenko,^{111,ll} A. Papageorgiou,¹¹¹ M. Pesaresi,¹¹¹ K. Petridis,¹¹¹ M. Pioppi,^{111,uu} D. M. Raymond,¹¹¹ S. Rogerson,¹¹¹ N. Rompotis,¹¹¹ A. Rose,¹¹¹ M. J. Ryan,¹¹¹ C. Seez,¹¹¹ P. Sharp,¹¹¹ A. Sparrow,¹¹¹ A. Tapper,¹¹¹ S. Tourneur,¹¹¹ M. Vazquez Acosta,¹¹¹ T. Virdee,¹¹¹ S. Wakefield,¹¹¹ N. Wardle,¹¹¹ D. Wardrope,¹¹¹ T. Whyntie,¹¹¹ M. Barrett,¹¹² M. Chadwick,¹¹² J. E. Cole,¹¹² P. R. Hobson,¹¹² A. Khan,¹¹² P. Kyberd,¹¹² D. Leslie,¹¹² W. Martin,¹¹² I. D. Reid,¹¹² L. Teodorescu,¹¹² M. Turner,¹¹² K. Hatakeyama,¹¹³ H. Liu,¹¹³ T. Scarborough,¹¹³ C. Henderson,¹¹⁴ A. Avetisyan,¹¹⁵ T. Bose,¹¹⁵ E. Carrera Jarrin,¹¹⁵ C. Fantasia,¹¹⁵ A. Heister,¹¹⁵ J. St. John,¹¹⁵ P. Lawson,¹¹⁵ D. Lazic,¹¹⁵ J. Rohlf,¹¹⁵ D. Sperka,¹¹⁵ L. Sulak,¹¹⁵ S. Bhattacharya,¹¹⁶ D. Cutts,¹¹⁶ A. Ferapontov,¹¹⁶ U. Heintz,¹¹⁶ S. Jabeen,¹¹⁶ G. Kukartsev,¹¹⁶ G. Landsberg,¹¹⁶ M. Luk,¹¹⁶ M. Narain,¹¹⁶ D. Nguyen,¹¹⁶ M. Segala,¹¹⁶ T. Sinthuprasith,¹¹⁶ T. Speer,¹¹⁶ K. V. Tsang,¹¹⁶ R. Breedon,¹¹⁷ G. Breto,¹¹⁷ M. Calderon De La Barca Sanchez,¹¹⁷ S. Chauhan,¹¹⁷ M. Chertok,¹¹⁷ J. Conway,¹¹⁷ R. Conway,¹¹⁷ P. T. Cox,¹¹⁷ J. Dolen,¹¹⁷ R. Erbacher,¹¹⁷ R. Houtz,¹¹⁷ W. Ko,¹¹⁷ A. Kopecky,¹¹⁷ R. Lander,¹¹⁷ O. Mall,¹¹⁷ T. Miceli,¹¹⁷ D. Pellett,¹¹⁷ J. Robles,¹¹⁷ B. Rutherford,¹¹⁷ M. Searle,¹¹⁷ J. Smith,¹¹⁷ M. Squires,¹¹⁷ M. Tripathi,¹¹⁷ R. Vasquez Sierra,¹¹⁷ V. Andreev,¹¹⁸ K. Arisaka,¹¹⁸ D. Cline,¹¹⁸ R. Cousins,¹¹⁸ J. Duris,¹¹⁸ S. Erhan,¹¹⁸ P. Everaerts,¹¹⁸ C. Farrell,¹¹⁸ J. Hauser,¹¹⁸ M. Ignatenko,¹¹⁸ C. Jarvis,¹¹⁸ C. Plager,¹¹⁸ G. Rakness,¹¹⁸ P. Schlein,^{118,a} J. Tucker,¹¹⁸ V. Valuev,¹¹⁸ M. Weber,¹¹⁸ J. Babb,¹¹⁹ R. Clare,¹¹⁹ J. Ellison,¹¹⁹ J. W. Gary,¹¹⁹ F. Giordano,¹¹⁹ G. Hanson,¹¹⁹ G. Y. Jeng,¹¹⁹ S. C. Kao,¹¹⁹ H. Liu,¹¹⁹ O. R. Long,¹¹⁹ A. Luthra,¹¹⁹ H. Nguyen,¹¹⁹ S. Paramesvaran,¹¹⁹ J. Sturdy,¹¹⁹ S. Sumowidagdo,¹¹⁹ R. Wilken,¹¹⁹ S. Wimpenny,¹¹⁹ W. Andrews,¹²⁰ J. G. Branson,¹²⁰ G. B. Cerati,¹²⁰ D. Evans,¹²⁰ F. Golf,¹²⁰ A. Holzner,¹²⁰ R. Kelley,¹²⁰ M. Lebourgeois,¹²⁰ J. Letts,¹²⁰ I. Macneill,¹²⁰ B. Mangano,¹²⁰ S. Padhi,¹²⁰ C. Palmer,¹²⁰ G. Petrucciani,¹²⁰ H. Pi,¹²⁰ M. Pieri,¹²⁰ R. Ranieri,¹²⁰ M. Sani,¹²⁰ I. Sfiligoi,¹²⁰ V. Sharma,¹²⁰ S. Simon,¹²⁰ E. Sudano,¹²⁰ M. Tadel,¹²⁰ Y. Tu,¹²⁰ A. Vartak,¹²⁰ S. Wasserbaech,^{120,vv} F. Würthwein,¹²⁰ A. Yagil,¹²⁰ J. Yoo,¹²⁰ D. Barge,¹²¹ R. Bellan,¹²¹ C. Campagnari,¹²¹ M. D'Alfonso,¹²¹ T. Danielson,¹²¹ K. Flowers,¹²¹ P. Geffert,¹²¹ C. George,¹²¹ J. Incandela,¹²¹ C. Justus,¹²¹ P. Kalavase,¹²¹ S. A. Koay,¹²¹ D. Kovalskyi,^{121,b} V. Krutelyov,¹²¹ S. Lowette,¹²¹ N. Mccoll,¹²¹ S. D. Mullin,¹²¹ V. Pavlunin,¹²¹ F. Rebassoo,¹²¹ J. Ribnik,¹²¹ J. Richman,¹²¹ R. Rossin,¹²¹ D. Stuart,¹²¹ W. To,¹²¹ J. R. Vlimant,¹²¹ C. West,¹²¹ A. Apresyan,¹²² A. Bornheim,¹²² J. Bunn,¹²² Y. Chen,¹²² E. Di Marco,¹²² J. Duarte,¹²² M. Gataullin,¹²² Y. Ma,¹²² A. Mott,¹²² H. B. Newman,¹²² C. Rogan,¹²² V. Timciuc,¹²² P. Traczyk,¹²² J. Veverka,¹²² R. Wilkinson,¹²² Y. Yang,¹²² R. Y. Zhu,¹²² B. Akgun,¹²³ R. Carroll,¹²³ T. Ferguson,¹²³ Y. Iiyama,¹²³ D. W. Jang,¹²³ S. Y. Jun,¹²³ Y. F. Liu,¹²³ M. Paulini,¹²³ J. Russ,¹²³ H. Vogel,¹²³ I. Vorobiev,¹²³ J. P. Cumalat,¹²⁴ M. E. Dinardo,¹²⁴ B. R. Drell,¹²⁴ C. J. Edelmaier,¹²⁴ W. T. Ford,¹²⁴ A. Gaz,¹²⁴ B. Heyburn,¹²⁴ E. Luiggi Lopez,¹²⁴ U. Nauenberg,¹²⁴ J. G. Smith,¹²⁴ K. Stenson,¹²⁴ K. A. Ulmer,¹²⁴ S. R. Wagner,¹²⁴ S. L. Zang,¹²⁴ L. Agostino,¹²⁵ J. Alexander,¹²⁵

A. Chatterjee,¹²⁵ N. Eggert,¹²⁵ L. K. Gibbons,¹²⁵ B. Heltsley,¹²⁵ W. Hopkins,¹²⁵ A. Khukhunaishvili,¹²⁵ B. Kreis,¹²⁵ G. Nicolas Kaufman,¹²⁵ J. R. Patterson,¹²⁵ D. Puigh,¹²⁵ A. Ryd,¹²⁵ E. Salvati,¹²⁵ X. Shi,¹²⁵ W. Sun,¹²⁵ W. D. Teo,¹²⁵ J. Thom,¹²⁵ J. Thompson,¹²⁵ J. Vaughan,¹²⁵ Y. Weng,¹²⁵ L. Winstrom,¹²⁵ P. Wittich,¹²⁵ A. Biselli,¹²⁶ G. Cirino,¹²⁶ D. Winn,¹²⁶ S. Abdullin,¹²⁷ M. Albrow,¹²⁷ J. Anderson,¹²⁷ G. Apollinari,¹²⁷ M. Atac,¹²⁷ J. A. Bakken,¹²⁷ L. A. T. Bauerdick,¹²⁷ A. Beretvas,¹²⁷ J. Berryhill,¹²⁷ P. C. Bhat,¹²⁷ I. Bloch,¹²⁷ K. Burkett,¹²⁷ J. N. Butler,¹²⁷ V. Chetluru,¹²⁷ H. W. K. Cheung,¹²⁷ F. Chlebana,¹²⁷ S. Cihangir,¹²⁷ W. Cooper,¹²⁷ D. P. Eartly,¹²⁷ V. D. Elvira,¹²⁷ S. Esen,¹²⁷ I. Fisk,¹²⁷ J. Freeman,¹²⁷ Y. Gao,¹²⁷ E. Gottschalk,¹²⁷ D. Green,¹²⁷ O. Gutsche,¹²⁷ J. Hanlon,¹²⁷ R. M. Harris,¹²⁷ J. Hirschauer,¹²⁷ B. Hooberman,¹²⁷ H. Jensen,¹²⁷ S. Jindariani,¹²⁷ M. Johnson,¹²⁷ U. Joshi,¹²⁷ B. Klima,¹²⁷ K. Kousouris,¹²⁷ S. Kunori,¹²⁷ S. Kwan,¹²⁷ C. Leonidopoulos,¹²⁷ D. Lincoln,¹²⁷ R. Lipton,¹²⁷ J. Lykken,¹²⁷ K. Maeshima,¹²⁷ J. M. Marraffino,¹²⁷ S. Maruyama,¹²⁷ D. Mason,¹²⁷ P. McBride,¹²⁷ T. Miao,¹²⁷ K. Mishra,¹²⁷ S. Mrenna,¹²⁷ Y. Musienko,^{127,ww} C. Newman-Holmes,¹²⁷ V. O'Dell,¹²⁷ J. Pivarski,¹²⁷ R. Pordes,¹²⁷ O. Prokofyev,¹²⁷ T. Schwarz,¹²⁷ E. Sexton-Kennedy,¹²⁷ S. Sharma,¹²⁷ W. J. Spalding,¹²⁷ L. Spiegel,¹²⁷ P. Tan,¹²⁷ L. Taylor,¹²⁷ S. Tkaczyk,¹²⁷ L. Uplegger,¹²⁷ E. W. Vaandering,¹²⁷ R. Vidal,¹²⁷ J. Whitmore,¹²⁷ W. Wu,¹²⁷ F. Yang,¹²⁷ F. Yumiceva,¹²⁷ J. C. Yun,¹²⁷ D. Acosta,¹²⁸ P. Avery,¹²⁸ D. Bourilkov,¹²⁸ M. Chen,¹²⁸ S. Das,¹²⁸ M. De Gruttola,¹²⁸ G. P. Di Giovanni,¹²⁸ D. Dobur,¹²⁸ A. Drozdetskiy,¹²⁸ R. D. Field,¹²⁸ M. Fisher,¹²⁸ Y. Fu,¹²⁸ I. K. Furic,¹²⁸ J. Gartner,¹²⁸ S. Goldberg,¹²⁸ J. Hugon,¹²⁸ B. Kim,¹²⁸ J. Konigsberg,¹²⁸ A. Korytov,¹²⁸ A. Kropivnitskaya,¹²⁸ T. Kypreos,¹²⁸ J. F. Low,¹²⁸ K. Matchev,¹²⁸ G. Mitselmakher,¹²⁸ L. Muniz,¹²⁸ M. Park,¹²⁸ R. Remington,¹²⁸ A. Rinkevicius,¹²⁸ M. Schmitt,¹²⁸ B. Scurlock,¹²⁸ P. Sellers,¹²⁸ N. Skhirtladze,¹²⁸ M. Snowball,¹²⁸ D. Wang,¹²⁸ J. Yelton,¹²⁸ M. Zakaria,¹²⁸ V. Gaultney,¹²⁹ L. M. Lebolo,¹²⁹ S. Linn,¹²⁹ P. Markowitz,¹²⁹ G. Martinez,¹²⁹ J. L. Rodriguez,¹²⁹ T. Adams,¹³⁰ A. Askew,¹³⁰ J. Bochenek,¹³⁰ J. Chen,¹³⁰ B. Diamond,¹³⁰ S. V. Gleyzer,¹³⁰ J. Haas,¹³⁰ S. Hagopian,¹³⁰ V. Hagopian,¹³⁰ M. Jenkins,¹³⁰ K. F. Johnson,¹³⁰ H. Prosper,¹³⁰ S. Sekmen,¹³⁰ V. Veeraraghavan,¹³⁰ M. M. Baarmand,¹³¹ B. Dorney,¹³¹ M. Hohlmann,¹³¹ H. Kalakhety,¹³¹ I. Vodopyanov,¹³¹ M. R. Adams,¹³² I. M. Anghel,¹³² L. Apanasevich,¹³² Y. Bai,¹³² V. E. Bazterra,¹³² R. R. Betts,¹³² J. Callner,¹³² R. Cavanaugh,¹³² C. Dragoiu,¹³² L. Gauthier,¹³² C. E. Gerber,¹³² D. J. Hofman,¹³² S. Khalatyan,¹³² G. J. Kunde,^{132,xx} F. Lacroix,¹³² M. Malek,¹³² C. O'Brien,¹³² C. Silkworth,¹³² C. Silvestre,¹³² D. Strom,¹³² N. Varelas,¹³² U. Akgun,¹³³ E. A. Albayrak,¹³³ B. Bilki,¹³³ W. Clarida,¹³³ F. Duru,¹³³ S. Griffiths,¹³³ C. K. Lae,¹³³ E. McCliment,¹³³ J.-P. Merlo,¹³³ H. Mermerkaya,^{133,yy} A. Mestvirishvili,¹³³ A. Moeller,¹³³ J. Nachtman,¹³³ C. R. Newsom,¹³³ E. Norbeck,¹³³ J. Olson,¹³³ Y. Onel,¹³³ F. Ozok,¹³³ S. Sen,¹³³ E. Tiras,¹³³ J. Wetzel,¹³³ T. Yetkin,¹³³ K. Yi,¹³³ B. A. Barnett,¹³⁴ B. Blumenfeld,¹³⁴ S. Bolognesi,¹³⁴ A. Bonato,¹³⁴ C. Eskew,¹³⁴ D. Fehling,¹³⁴ G. Giurgiu,¹³⁴ A. V. Gritsan,¹³⁴ Z. J. Guo,¹³⁴ G. Hu,¹³⁴ P. Maksimovic,¹³⁴ S. Rappoccio,¹³⁴ M. Swartz,¹³⁴ N. V. Tran,¹³⁴ A. Whitbeck,¹³⁴ P. Baringer,¹³⁵ A. Bean,¹³⁵ G. Benelli,¹³⁵ O. Grachov,¹³⁵ R. P. Kenny Iii,¹³⁵ M. Murray,¹³⁵ D. Noonan,¹³⁵ S. Sanders,¹³⁵ R. Stringer,¹³⁵ G. Tinti,¹³⁵ J. S. Wood,¹³⁵ V. Zhukova,¹³⁵ A. F. Barfuss,¹³⁶ T. Bolton,¹³⁶ I. Chakaberia,¹³⁶ A. Ivanov,¹³⁶ S. Khalil,¹³⁶ M. Makouski,¹³⁶ Y. Maravin,¹³⁶ S. Shrestha,¹³⁶ I. Svintradze,¹³⁶ J. Gronberg,¹³⁷ D. Lange,¹³⁷ D. Wright,¹³⁷ A. Baden,¹³⁸ M. Boutemour,¹³⁸ B. Calvert,¹³⁸ S. C. Eno,¹³⁸ J. A. Gomez,¹³⁸ N. J. Hadley,¹³⁸ R. G. Kellogg,¹³⁸ M. Kirn,¹³⁸ Y. Lu,¹³⁸ A. C. Mignerey,¹³⁸ A. Peterman,¹³⁸ K. Rossato,¹³⁸ P. Rumerio,¹³⁸ A. Skuja,¹³⁸ J. Temple,¹³⁸ M. B. Tonjes,¹³⁸ S. C. Tonwar,¹³⁸ E. Twedt,¹³⁸ B. Alver,¹³⁹ G. Bauer,¹³⁹ J. Bendavid,¹³⁹ W. Busza,¹³⁹ E. Butz,¹³⁹ I. A. Cali,¹³⁹ M. Chan,¹³⁹ V. Dutta,¹³⁹ G. Gomez Ceballos,¹³⁹ M. Goncharov,¹³⁹ K. A. Hahn,¹³⁹ P. Harris,¹³⁹ Y. Kim,¹³⁹ M. Klute,¹³⁹ Y.-J. Lee,¹³⁹ W. Li,¹³⁹ P. D. Luckey,¹³⁹ T. Ma,¹³⁹ S. Nahn,¹³⁹ C. Paus,¹³⁹ D. Ralph,¹³⁹ C. Roland,¹³⁹ G. Roland,¹³⁹ M. Rudolph,¹³⁹ G. S. F. Stephens,¹³⁹ F. Stöckli,¹³⁹ K. Sumorok,¹³⁹ K. Sung,¹³⁹ D. Velicanu,¹³⁹ E. A. Wenger,¹³⁹ R. Wolf,¹³⁹ B. Wyslouch,¹³⁹ S. Xie,¹³⁹ M. Yang,¹³⁹ Y. Yilmaz,¹³⁹ A. S. Yoon,¹³⁹ M. Zanetti,¹³⁹ S. I. Cooper,¹⁴⁰ P. Cushman,¹⁴⁰ B. Dahmes,¹⁴⁰ A. De Benedetti,¹⁴⁰ G. Franzoni,¹⁴⁰ A. Gude,¹⁴⁰ J. Haupt,¹⁴⁰ K. Klapoetke,¹⁴⁰ Y. Kubota,¹⁴⁰ J. Mans,¹⁴⁰ N. Pastika,¹⁴⁰ V. Rekovic,¹⁴⁰ R. Rusack,¹⁴⁰ M. Sasseville,¹⁴⁰ A. Singovsky,¹⁴⁰ N. Tambe,¹⁴⁰ J. Turkewitz,¹⁴⁰ L. M. Cremaldi,¹⁴¹ R. Godang,¹⁴¹ R. Kroeger,¹⁴¹ L. Perera,¹⁴¹ R. Rahmat,¹⁴¹ D. A. Sanders,¹⁴¹ D. Summers,¹⁴¹ E. Avdeeva,¹⁴² K. Bloom,¹⁴² S. Bose,¹⁴² J. Butt,¹⁴² D. R. Claes,¹⁴² A. Dominguez,¹⁴² M. Eads,¹⁴² P. Jindal,¹⁴² J. Keller,¹⁴² I. Kravchenko,¹⁴² J. Lazo-Flores,¹⁴² H. Malbouisson,¹⁴² S. Malik,¹⁴² G. R. Snow,¹⁴² U. Baur,¹⁴³ A. Godshalk,¹⁴³ I. Iashvili,¹⁴³ S. Jain,¹⁴³ A. Kharchilava,¹⁴³ A. Kumar,¹⁴³ K. Smith,¹⁴³ Z. Wan,¹⁴³ G. Alverson,¹⁴⁴ E. Barberis,¹⁴⁴ D. Baumgartel,¹⁴⁴ M. Chasco,¹⁴⁴ D. Trocino,¹⁴⁴ D. Wood,¹⁴⁴ J. Zhang,¹⁴⁴ A. Anastassov,¹⁴⁵ A. Kubik,¹⁴⁵ N. Mucia,¹⁴⁵ N. Odell,¹⁴⁵ R. A. Oforzynski,¹⁴⁵ B. Pollack,¹⁴⁵ A. Pozdnyakov,¹⁴⁵ M. Schmitt,¹⁴⁵ S. Stoynev,¹⁴⁵ M. Velasco,¹⁴⁵ S. Won,¹⁴⁵ L. Antonelli,¹⁴⁶ D. Berry,¹⁴⁶ A. Brinkerhoff,¹⁴⁶ M. Hildreth,¹⁴⁶ C. Jessop,¹⁴⁶ D. J. Karmgard,¹⁴⁶ J. Kolb,¹⁴⁶ T. Kolberg,¹⁴⁶ K. Lannon,¹⁴⁶ W. Luo,¹⁴⁶ S. Lynch,¹⁴⁶ N. Marinelli,¹⁴⁶

D. M. Morse,¹⁴⁶ T. Pearson,¹⁴⁶ R. Ruchti,¹⁴⁶ J. Slaunwhite,¹⁴⁶ N. Valls,¹⁴⁶ M. Wayne,¹⁴⁶ J. Ziegler,¹⁴⁶ B. Bylsma,¹⁴⁷ L. S. Durkin,¹⁴⁷ C. Hill,¹⁴⁷ P. Killewald,¹⁴⁷ K. Kotov,¹⁴⁷ T. Y. Ling,¹⁴⁷ M. Rodenburg,¹⁴⁷ C. Vuosalo,¹⁴⁷ G. Williams,¹⁴⁷ N. Adam,¹⁴⁸ E. Berry,¹⁴⁸ P. Elmer,¹⁴⁸ D. Gerbaudo,¹⁴⁸ V. Halyo,¹⁴⁸ P. Hebda,¹⁴⁸ A. Hunt,¹⁴⁸ E. Laird,¹⁴⁸ D. Lopes Pegna,¹⁴⁸ P. Lujan,¹⁴⁸ D. Marlow,¹⁴⁸ T. Medvedeva,¹⁴⁸ M. Mooney,¹⁴⁸ J. Olsen,¹⁴⁸ P. Piroué,¹⁴⁸ X. Quan,¹⁴⁸ A. Raval,¹⁴⁸ H. Saka,¹⁴⁸ D. Stickland,¹⁴⁸ C. Tully,¹⁴⁸ J. S. Werner,¹⁴⁸ A. Zuranski,¹⁴⁸ J. G. Acosta,¹⁴⁹ X. T. Huang,¹⁴⁹ A. Lopez,¹⁴⁹ H. Mendez,¹⁴⁹ S. Oliveros,¹⁴⁹ J. E. Ramirez Vargas,¹⁴⁹ A. Zatserklyaniy,¹⁴⁹ E. Alagoz,¹⁵⁰ V. E. Barnes,¹⁵⁰ D. Benedetti,¹⁵⁰ G. Bolla,¹⁵⁰ L. Borrello,¹⁵⁰ D. Bortoletto,¹⁵⁰ M. De Mattia,¹⁵⁰ A. Everett,¹⁵⁰ L. Gutay,¹⁵⁰ Z. Hu,¹⁵⁰ M. Jones,¹⁵⁰ O. Koybasi,¹⁵⁰ M. Kress,¹⁵⁰ A. T. Laasanen,¹⁵⁰ N. Leonardo,¹⁵⁰ V. Maroussov,¹⁵⁰ P. Merkel,¹⁵⁰ D. H. Miller,¹⁵⁰ N. Neumeister,¹⁵⁰ I. Shipsey,¹⁵⁰ D. Silvers,¹⁵⁰ A. Svyatkovskiy,¹⁵⁰ M. Vidal Marono,¹⁵⁰ H. D. Yoo,¹⁵⁰ J. Zablocki,¹⁵⁰ Y. Zheng,¹⁵⁰ S. Guragain,¹⁵¹ N. Parashar,¹⁵¹ A. Adair,¹⁵² C. Boulahouache,¹⁵² V. Cuplov,¹⁵² K. M. Ecklund,¹⁵² F. J. M. Geurts,¹⁵² B. P. Padley,¹⁵² R. Redjimi,¹⁵² J. Roberts,¹⁵² J. Zabel,¹⁵² B. Betchart,¹⁵³ A. Bodek,¹⁵³ Y. S. Chung,¹⁵³ R. Covarelli,¹⁵³ P. de Barbaro,¹⁵³ R. Demina,¹⁵³ Y. Eshaq,¹⁵³ H. Flacher,¹⁵³ A. Garcia-Bellido,¹⁵³ P. Goldenzweig,¹⁵³ Y. Gotra,¹⁵³ J. Han,¹⁵³ A. Harel,¹⁵³ D. C. Miner,¹⁵³ G. Petrillo,¹⁵³ W. Sakumoto,¹⁵³ D. Vishnevskiy,¹⁵³ M. Zielinski,¹⁵³ A. Bhatti,¹⁵⁴ R. Ciesielski,¹⁵⁴ L. Demortier,¹⁵⁴ K. Goulianos,¹⁵⁴ G. Lungu,¹⁵⁴ S. Malik,¹⁵⁴ C. Mesropian,¹⁵⁴ S. Arora,¹⁵⁵ O. Atramentov,¹⁵⁵ A. Barker,¹⁵⁵ J. P. Chou,¹⁵⁵ C. Contreras-Campana,¹⁵⁵ E. Contreras-Campana,¹⁵⁵ D. Duggan,¹⁵⁵ D. Ferencek,¹⁵⁵ Y. Gershtein,¹⁵⁵ R. Gray,¹⁵⁵ E. Halkiadakis,¹⁵⁵ D. Hidas,¹⁵⁵ D. Hits,¹⁵⁵ A. Lath,¹⁵⁵ S. Panwalkar,¹⁵⁵ M. Park,¹⁵⁵ R. Patel,¹⁵⁵ A. Richards,¹⁵⁵ K. Rose,¹⁵⁵ S. Salur,¹⁵⁵ S. Schnetzer,¹⁵⁵ S. Somalwar,¹⁵⁵ R. Stone,¹⁵⁵ S. Thomas,¹⁵⁵ G. Cerizza,¹⁵⁶ M. Hollingsworth,¹⁵⁶ S. Spanier,¹⁵⁶ Z. C. Yang,¹⁵⁶ A. York,¹⁵⁶ R. Eusebi,¹⁵⁷ W. Flanagan,¹⁵⁷ J. Gilmore,¹⁵⁷ T. Kamon,^{157,zz} V. Khotilovich,¹⁵⁷ R. Montalvo,¹⁵⁷ I. Osipenkov,¹⁵⁷ Y. Pakhotin,¹⁵⁷ A. Perloff,¹⁵⁷ J. Roe,¹⁵⁷ A. Safonov,¹⁵⁷ S. Sengupta,¹⁵⁷ I. Suarez,¹⁵⁷ A. Tatarinov,¹⁵⁷ D. Toback,¹⁵⁷ N. Akchurin,¹⁵⁸ C. Bardak,¹⁵⁸ J. Damgov,¹⁵⁸ P. R. Duerdo,¹⁵⁸ C. Jeong,¹⁵⁸ K. Kovitanggoon,¹⁵⁸ S. W. Lee,¹⁵⁸ T. Libeiro,¹⁵⁸ P. Mane,¹⁵⁸ Y. Roh,¹⁵⁸ A. Sill,¹⁵⁸ I. Volobouev,¹⁵⁸ R. Wigmans,¹⁵⁸ E. Yazgan,¹⁵⁸ E. Appelt,¹⁵⁹ E. Brownson,¹⁵⁹ D. Engh,¹⁵⁹ C. Florez,¹⁵⁹ W. Gabella,¹⁵⁹ A. Gurrola,¹⁵⁹ M. Issah,¹⁵⁹ W. Johns,¹⁵⁹ C. Johnston,¹⁵⁹ P. Kurt,¹⁵⁹ C. Maguire,¹⁵⁹ A. Melo,¹⁵⁹ P. Sheldon,¹⁵⁹ B. Snook,¹⁵⁹ S. Tuo,¹⁵⁹ J. Velkovska,¹⁵⁹ M. W. Arenton,¹⁶⁰ M. Balazs,¹⁶⁰ S. Boutle,¹⁶⁰ S. Conetti,¹⁶⁰ B. Cox,¹⁶⁰ B. Francis,¹⁶⁰ S. Goadhouse,¹⁶⁰ J. Goodell,¹⁶⁰ R. Hirosky,¹⁶⁰ A. Ledovskoy,¹⁶⁰ C. Lin,¹⁶⁰ C. Neu,¹⁶⁰ J. Wood,¹⁶⁰ R. Yohay,¹⁶⁰ S. Gollapinni,¹⁶¹ R. Harr,¹⁶¹ P. E. Karchin,¹⁶¹ C. Kottachchi Kankanamge Don,¹⁶¹ P. Lamichhane,¹⁶¹ M. Mattson,¹⁶¹ C. Milstène,¹⁶¹ A. Sakharov,¹⁶¹ M. Anderson,¹⁶² M. Bachtis,¹⁶² D. Belknap,¹⁶² J. N. Bellinger,¹⁶² J. Bernardini,¹⁶² D. Carlsmith,¹⁶² M. Cepeda,¹⁶² S. Dasu,¹⁶² J. Efron,¹⁶² E. Friis,¹⁶² L. Gray,¹⁶² K. S. Grogg,¹⁶² M. Grothe,¹⁶² R. Hall-Wilton,¹⁶² M. Herndon,¹⁶² A. Hervé,¹⁶² P. Klabbers,¹⁶² J. Klukas,¹⁶² A. Lanaro,¹⁶² C. Lazaridis,¹⁶² J. Leonard,¹⁶² R. Loveless,¹⁶² A. Mohapatra,¹⁶² I. Ojalvo,¹⁶² G. A. Pierro,¹⁶² I. Ross,¹⁶² A. Savin,¹⁶² W. H. Smith,¹⁶² J. Swanson,¹⁶² and M. Weinberg¹⁶²

(CMS Collaboration)

¹*Yerevan Physics Institute, Yerevan, Armenia*²*Institut für Hochenergiephysik der OeAW, Wien, Austria*³*National Centre for Particle and High Energy Physics, Minsk, Belarus*⁴*Universiteit Antwerpen, Antwerpen, Belgium*⁵*Vrije Universiteit Brussel, Brussel, Belgium*⁶*Université Libre de Bruxelles, Bruxelles, Belgium*⁷*Ghent University, Ghent, Belgium*⁸*Université Catholique de Louvain, Louvain-la-Neuve, Belgium*⁹*Université de Mons, Mons, Belgium*¹⁰*Centro Brasileiro de Pesquisas Físicas, Rio de Janeiro, Brazil*¹¹*Universidade do Estado do Rio de Janeiro, Rio de Janeiro, Brazil*¹²*Instituto de Física Teórica, Universidade Estadual Paulista, Sao Paulo, Brazil*¹³*Institute for Nuclear Research and Nuclear Energy, Sofia, Bulgaria*¹⁴*University of Sofia, Sofia, Bulgaria*¹⁵*Institute of High Energy Physics, Beijing, China*¹⁶*State Key Laboratory of Nuclear Physics and Technology, Peking University, Beijing, China*¹⁷*Universidad de Los Andes, Bogota, Colombia*¹⁸*Technical University of Split, Split, Croatia*¹⁹*University of Split, Split, Croatia*

- ²⁰*Institute Rudjer Boskovic, Zagreb, Croatia*
- ²¹*University of Cyprus, Nicosia, Cyprus*
- ²²*Charles University, Prague, Czech Republic*
- ²³*Academy of Scientific Research and Technology of the Arab Republic of Egypt, Egyptian Network of High Energy Physics, Cairo, Egypt*
- ²⁴*National Institute of Chemical Physics and Biophysics, Tallinn, Estonia*
- ²⁵*Department of Physics, University of Helsinki, Helsinki, Finland*
- ²⁶*Helsinki Institute of Physics, Helsinki, Finland*
- ²⁷*Lappeenranta University of Technology, Lappeenranta, Finland*
- ²⁸*Laboratoire d'Annecy-le-Vieux de Physique des Particules, IN2P3-CNRS, Annecy-le-Vieux, France*
- ²⁹*DSM/IRFU, CEA/Saclay, Gif-sur-Yvette, France*
- ³⁰*Laboratoire Leprince-Ringuet, Ecole Polytechnique, IN2P3-CNRS, Palaiseau, France*
- ³¹*Institut Pluridisciplinaire Hubert Curien, Université de Strasbourg, Université de Haute Alsace Mulhouse, CNRS/IN2P3, Strasbourg, France*
- ³²*Centre de Calcul de l'Institut National de Physique Nucléaire et de Physique des Particules (IN2P3), Villeurbanne, France*
- ³³*Université de Lyon, Université Claude Bernard Lyon 1, CNRS-IN2P3, Institut de Physique Nucléaire de Lyon, Villeurbanne, France*
- ³⁴*Institute of High Energy Physics and Informatization, Tbilisi State University, Tbilisi, Georgia*
- ³⁵*RWTH Aachen University, I. Physikalisches Institut, Aachen, Germany*
- ³⁶*RWTH Aachen University, III. Physikalisches Institut A, Aachen, Germany*
- ³⁷*RWTH Aachen University, III. Physikalisches Institut B, Aachen, Germany*
- ³⁸*Deutsches Elektronen-Synchrotron, Hamburg, Germany*
- ³⁹*University of Hamburg, Hamburg, Germany*
- ⁴⁰*Institut für Experimentelle Kernphysik, Karlsruhe, Germany*
- ⁴¹*Institute of Nuclear Physics "Demokritos," Aghia Paraskevi, Greece*
- ⁴²*University of Athens, Athens, Greece*
- ⁴³*University of Ioánnina, Ioánnina, Greece*
- ⁴⁴*KFKI Research Institute for Particle and Nuclear Physics, Budapest, Hungary*
- ⁴⁵*Institute of Nuclear Research ATOMKI, Debrecen, Hungary*
- ⁴⁶*University of Debrecen, Debrecen, Hungary*
- ⁴⁷*Panjab University, Chandigarh, India*
- ⁴⁸*University of Delhi, Delhi, India*
- ⁴⁹*Saha Institute of Nuclear Physics, Kolkata, India*
- ⁵⁰*Bhabha Atomic Research Centre, Mumbai, India*
- ⁵¹*Tata Institute of Fundamental Research–EHEP, Mumbai, India*
- ⁵²*Tata Institute of Fundamental Research–HECR, Mumbai, India*
- ⁵³*Institute for Research and Fundamental Sciences (IPM), Tehran, Iran*
- ^{54a}*INFN Sezione di Bari, Bari, Italy*
- ^{54b}*Università di Bari, Bari, Italy*
- ^{54c}*Politecnico di Bari, Bari, Italy*
- ^{55a}*INFN Sezione di Bologna, Bologna, Italy*
- ^{55b}*Università di Bologna, Bologna, Italy*
- ^{56a}*INFN Sezione di Catania, Catania, Italy*
- ^{56b}*Università di Catania, Catania, Italy*
- ^{57a}*INFN Sezione di Firenze, Firenze, Italy*
- ^{57b}*Università di Firenze, Firenze, Italy*
- ⁵⁸*INFN Laboratori Nazionali di Frascati, Frascati, Italy*
- ⁵⁹*INFN Sezione di Genova, Genova, Italy*
- ^{60a}*INFN Sezione di Milano-Bicocca, Milano, Italy*
- ^{60b}*Università di Milano-Bicocca, Milano, Italy*
- ^{61a}*INFN Sezione di Napoli, Napoli, Italy*
- ^{61b}*Università di Napoli "Federico II," Napoli, Italy*
- ^{62a}*INFN Sezione di Padova, Padova, Italy*
- ^{62b}*Università di Padova, Padova, Italy*
- ^{62c}*Università di Trento (Trento), Padova, Italy*
- ^{63a}*INFN Sezione di Pavia, Pavia, Italy*
- ^{63b}*Università di Pavia, Pavia, Italy*
- ^{64a}*INFN Sezione di Perugia, Italy*
- ^{64b}*Università di Perugia, Italy*
- ^{65a}*INFN Sezione di Pisa, Pisa, Italy*
- ^{65b}*Università di Pisa, Pisa, Italy*
- ^{65c}*Scuola Normale Superiore di Pisa, Pisa, Italy*

- ^{66a}INFN Sezione di Roma, Roma, Italy
^{66b}Università di Roma “La Sapienza,” Roma, Italy
^{67a}INFN Sezione di Torino, Torino, Italy
^{67b}Università di Torino, Torino, Italy
^{67c}Università del Piemonte Orientale (Novara), Torino, Italy
^{68a}INFN Sezione di Trieste, Trieste, Italy
^{68b}Università di Trieste, Trieste, Italy
⁶⁹Kangwon National University, Chunchon, Korea
⁷⁰Kyungpook National University, Daegu, Korea
⁷¹Chonnam National University, Institute for Universe and Elementary Particles, Kwangju, Korea
⁷²Konkuk University, Seoul, Korea
⁷³Korea University, Seoul, Korea
⁷⁴University of Seoul, Seoul, Korea
⁷⁵Sungkyunkwan University, Suwon, Korea
⁷⁶Vilnius University, Vilnius, Lithuania
⁷⁷Centro de Investigacion y de Estudios Avanzados del IPN, Mexico City, Mexico
⁷⁸Universidad Iberoamericana, Mexico City, Mexico
⁷⁹Benemerita Universidad Autonoma de Puebla, Puebla, Mexico
⁸⁰Universidad Autónoma de San Luis Potosí, San Luis Potosí, Mexico
⁸¹University of Auckland, Auckland, New Zealand
⁸²University of Canterbury, Christchurch, New Zealand
⁸³National Centre for Physics, Quaid-I-Azam University, Islamabad, Pakistan
⁸⁴Institute of Experimental Physics, Faculty of Physics, University of Warsaw, Warsaw, Poland
⁸⁵Soltan Institute for Nuclear Studies, Warsaw, Poland
⁸⁶Laboratório de Instrumentação e Física Experimental de Partículas, Lisboa, Portugal
⁸⁷Joint Institute for Nuclear Research, Dubna, Russia
⁸⁸Petersburg Nuclear Physics Institute, Gatchina (St Petersburg), Russia
⁸⁹Institute for Nuclear Research, Moscow, Russia
⁹⁰Institute for Theoretical and Experimental Physics, Moscow, Russia
⁹¹Moscow State University, Moscow, Russia
⁹²P. N. Lebedev Physical Institute, Moscow, Russia
⁹³State Research Center of Russian Federation, Institute for High Energy Physics, Protvino, Russia
⁹⁴University of Belgrade, Faculty of Physics and Vinca Institute of Nuclear Sciences, Belgrade, Serbia
⁹⁵Centro de Investigaciones Energéticas Medioambientales y Tecnológicas (CIEMAT), Madrid, Spain
⁹⁶Universidad Autónoma de Madrid, Madrid, Spain
⁹⁷Universidad de Oviedo, Oviedo, Spain
⁹⁸Instituto de Física de Cantabria (IFCA), CSIC-Universidad de Cantabria, Santander, Spain
⁹⁹CERN, European Organization for Nuclear Research, Geneva, Switzerland
¹⁰⁰Paul Scherrer Institut, Villigen, Switzerland
¹⁰¹Institute for Particle Physics, ETH Zurich, Zurich, Switzerland
¹⁰²Universität Zürich, Zurich, Switzerland
¹⁰³National Central University, Chung-Li, Taiwan
¹⁰⁴National Taiwan University (NTU), Taipei, Taiwan
¹⁰⁵Cukurova University, Adana, Turkey
¹⁰⁶Middle East Technical University, Physics Department, Ankara, Turkey
¹⁰⁷Bogazici University, Istanbul, Turkey
¹⁰⁸National Scientific Center, Kharkov Institute of Physics and Technology, Kharkov, Ukraine
¹⁰⁹University of Bristol, Bristol, United Kingdom
¹¹⁰Rutherford Appleton Laboratory, Didcot, United Kingdom
¹¹¹Imperial College, London, United Kingdom
¹¹²Brunel University, Uxbridge, United Kingdom
¹¹³Baylor University, Waco, Texas, USA
¹¹⁴The University of Alabama, Tuscaloosa, Alabama, USA
¹¹⁵Boston University, Boston, Massachusetts, USA
¹¹⁶Brown University, Providence, Rhode Island, USA
¹¹⁷University of California, Davis, Davis, California, USA
¹¹⁸University of California, Los Angeles, Los Angeles, California, USA
¹¹⁹University of California, Riverside, Riverside, California, USA
¹²⁰University of California, San Diego, La Jolla, California, USA
¹²¹University of California, Santa Barbara, Santa Barbara, California, USA
¹²²California Institute of Technology, Pasadena, California, USA

- ¹²³*Carnegie Mellon University, Pittsburgh, Pennsylvania, USA*
¹²⁴*University of Colorado at Boulder, Boulder, Colorado, USA*
¹²⁵*Cornell University, Ithaca, New York, USA*
¹²⁶*Fairfield University, Fairfield, Connecticut, USA*
¹²⁷*Fermi National Accelerator Laboratory, Batavia, Illinois, USA*
¹²⁸*University of Florida, Gainesville, Florida, USA*
¹²⁹*Florida International University, Miami, Florida, USA*
¹³⁰*Florida State University, Tallahassee, Florida, USA*
¹³¹*Florida Institute of Technology, Melbourne, Florida, USA*
¹³²*University of Illinois at Chicago (UIC), Chicago, Illinois, USA*
¹³³*The University of Iowa, Iowa City, Iowa, USA*
¹³⁴*Johns Hopkins University, Baltimore, Maryland, USA*
¹³⁵*The University of Kansas, Lawrence, Kansas, USA*
¹³⁶*Kansas State University, Manhattan, Kansas, USA*
¹³⁷*Lawrence Livermore National Laboratory, Livermore, California, USA*
¹³⁸*University of Maryland, College Park, Maryland, USA*
¹³⁹*Massachusetts Institute of Technology, Cambridge, Massachusetts, USA*
¹⁴⁰*University of Minnesota, Minneapolis, Minnesota, USA*
¹⁴¹*University of Mississippi, University, Mississippi, USA*
¹⁴²*University of Nebraska-Lincoln, Lincoln, Nebraska, USA*
¹⁴³*State University of New York at Buffalo, Buffalo, New York, USA*
¹⁴⁴*Northeastern University, Boston, Massachusetts, USA*
¹⁴⁵*Northwestern University, Evanston, Illinois, USA*
¹⁴⁶*University of Notre Dame, Notre Dame, Indiana, USA*
¹⁴⁷*The Ohio State University, Columbus, Ohio, USA*
¹⁴⁸*Princeton University, Princeton, New Jersey, USA*
¹⁴⁹*University of Puerto Rico, Mayaguez, Puerto Rico*
¹⁵⁰*Purdue University, West Lafayette, Indiana, USA*
¹⁵¹*Purdue University Calumet, Hammond, Indiana, USA*
¹⁵²*Rice University, Houston, Texas, USA*
¹⁵³*University of Rochester, Rochester, New York, USA*
¹⁵⁴*The Rockefeller University, New York, New York, USA*
¹⁵⁵*Rutgers, the State University of New Jersey, Piscataway, New Jersey, USA*
¹⁵⁶*University of Tennessee, Knoxville, Tennessee, USA*
¹⁵⁷*Texas A&M University, College Station, Texas, USA*
¹⁵⁸*Texas Tech University, Lubbock, Texas, USA*
¹⁵⁹*Vanderbilt University, Nashville, Tennessee, USA*
¹⁶⁰*University of Virginia, Charlottesville, Virginia, USA*
¹⁶¹*Wayne State University, Detroit, Michigan, USA*
¹⁶²*University of Wisconsin, Madison, Wisconsin, USA*

^aDeceased.

^bAlso at CERN, European Organization for Nuclear Research, Geneva, Switzerland.

^cAlso at National Institute of Chemical Physics and Biophysics, Tallinn, Estonia.

^dAlso at Universidade Federal do ABC, Santo Andre, Brazil.

^eAlso at California Institute of Technology, Pasadena, California, USA.

^fAlso at Laboratoire Leprince-Ringuet, Ecole Polytechnique, IN2P3-CNRS, Palaiseau, France.

^gAlso at Suez Canal University, Suez, Egypt.

^hAlso at Cairo University, Cairo, Egypt.

ⁱAlso at British University, Cairo, Egypt.

^jAlso at Fayoum University, El-Fayoum, Egypt.

^kAlso at Ain Shams University, Cairo, Egypt.

^lAlso at Soltan Institute for Nuclear Studies, Warsaw, Poland.

^mAlso at Université de Haute-Alsace, Mulhouse, France.

ⁿAlso at Moscow State University, Moscow, Russia.

^oAlso at Brandenburg University of Technology, Cottbus, Germany.

^pAlso at Institute of Nuclear Research ATOMKI, Debrecen, Hungary.

^qAlso at Eötvös Loránd University, Budapest, Hungary.

^rAlso at Tata Institute of Fundamental Research-HECR, Mumbai, India.

- ^sAlso at University of Visva-Bharati, Santiniketan, India.
- ^tAlso at Sharif University of Technology, Tehran, Iran.
- ^uAlso at Isfahan University of Technology, Isfahan, Iran.
- ^vAlso at Shiraz University, Shiraz, Iran.
- ^wAlso at Plasma Physics Research Center, Science and Research Branch, Islamic Azad University, Teheran, Iran.
- ^xAlso at Facoltà Ingegneria Università di Roma, Roma, Italy.
- ^yAlso at Università della Basilicata, Potenza, Italy.
- ^zAlso at Laboratori Nazionali di Legnaro dell' INFN, Legnaro, Italy.
- ^{aa}Also at Università degli studi di Siena, Siena, Italy.
- ^{bb}Also at Faculty of Physics of University of Belgrade, Belgrade, Serbia.
- ^{cc}Also at University of California, Los Angeles, Los Angeles, California, USA.
- ^{dd}Also at University of Florida, Gainesville, Florida, USA.
- ^{ee}Also at Scuola Normale e Sezione dell' INFN, Pisa, Italy.
- ^{ff}Also at INFN Sezione di Roma, Università di Roma "La Sapienza," Roma, Italy.
- ^{gg}Also at University of Athens, Athens, Greece.
- ^{hh}Also at Rutherford Appleton Laboratory, Didcot, United Kingdom.
- ⁱⁱAlso at The University of Kansas, Lawrence, Kansas, USA.
- ^{jj}Also at University of Belgrade, Faculty of Physics and Vinca Institute of Nuclear Sciences, Belgrade, Serbia.
- ^{kk}Also at Paul Scherrer Institut, Villigen, Switzerland.
- ^{ll}Also at Institute for Theoretical and Experimental Physics, Moscow, Russia.
- ^{mm}Also at Gaziosmanpasa University, Tokat, Turkey.
- ⁿⁿAlso at Adiyaman University, Adiyaman, Turkey.
- ^{oo}Also at The University of Iowa, Iowa City, Iowa, USA.
- ^{pp}Also at Mersin University, Mersin, Turkey.
- ^{qq}Also at Kafkas University, Kars, Turkey.
- ^{rr}Also at Suleyman Demirel University, Isparta, Turkey.
- ^{ss}Also at Ege University, Izmir, Turkey.
- ^{tt}Also at School of Physics and Astronomy, University of Southampton, Southampton, United Kingdom.
- ^{uu}Also at INFN Sezione di Perugia, Università di Perugia, Perugia, Italy.
- ^{vv}Also at Utah Valley University, Orem, Utah, USA.
- ^{ww}Also at Institute for Nuclear Research, Moscow, Russia.
- ^{xx}Also at Los Alamos National Laboratory, Los Alamos, New Mexico, USA.
- ^{yy}Also at Erzincan University, Erzincan, Turkey.
- ^{zz}Also at Kyungpook National University, Daegu, Korea.